

#### Magnetosheath jets and their effects under different IMF orientations



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#### Magnetosheath jets: enhancements of $P_{\rm dyn} = \rho v^2$

- A jet triggering magnetopause reconnection (Hietala et al., 2018)
- Jets linked to triggering tail reconnection and a substorm during northward IMF (Nykyri et al., 2019)

Could jets statistically affect the occurrence of magnetic reconnection at the subsolar magnetopause?

#### **Statistical study:**

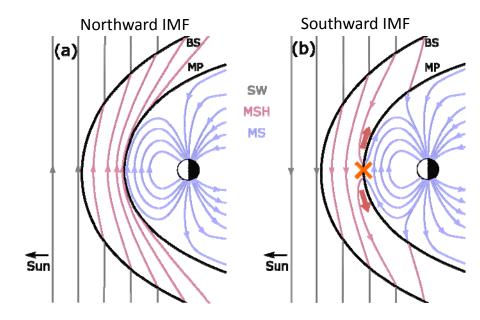
- 2,736.9 hours of THEMIS 2008—2011 data from the subsolar magnetosheath (Angelopoulos, 2008)
- 2,859 jets with Plaschke et al. (2013) criterion of earthward  $P_{\rm dyn}$  going above ½ SW  $P_{\rm dyn}$
- OMNI solar wind data (King and Papitashvili, 2005)

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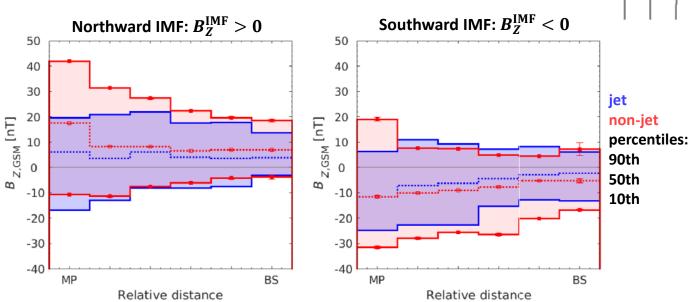
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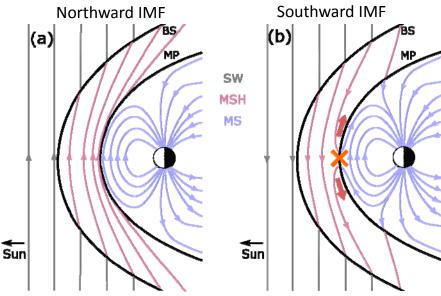
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- **GSM**  $B_Z$  is the most important parameter for reconnection in this region
  - ightarrow Comparison of  $B_Z$  within jets and ambient magnetosheath samples during similar IMF cone angle conditions

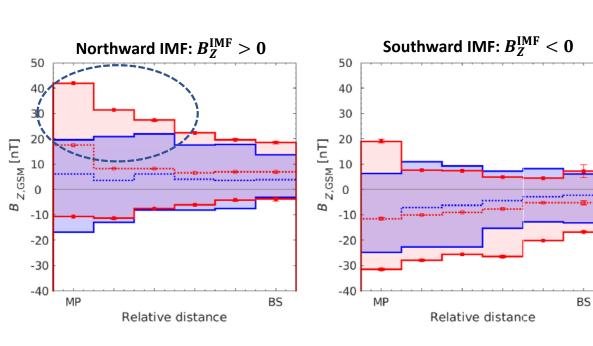


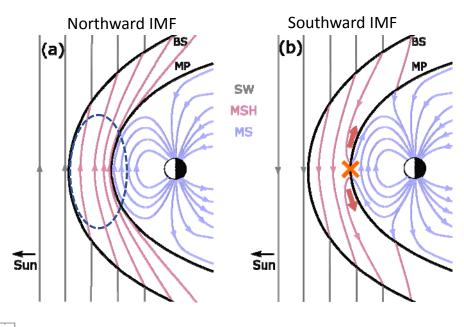
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jet

90th 50th

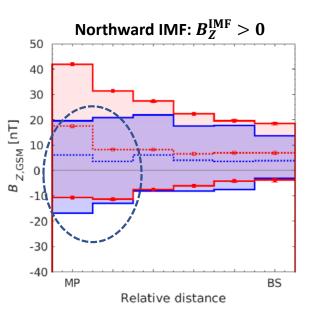
10th

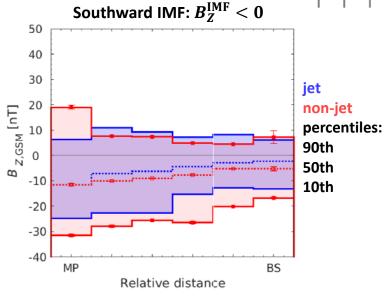
non-jet

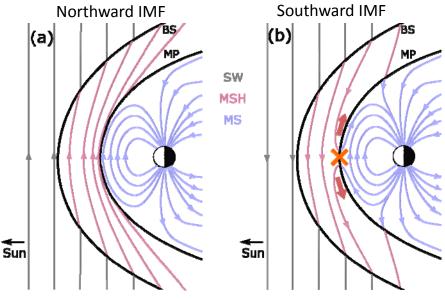
percentiles:

Magnetic pile-up during northward IMF

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  - $\rightarrow$  Comparison of  $B_Z$  within jets and ambient magnetosheath samples during similar IMF cone angle conditions



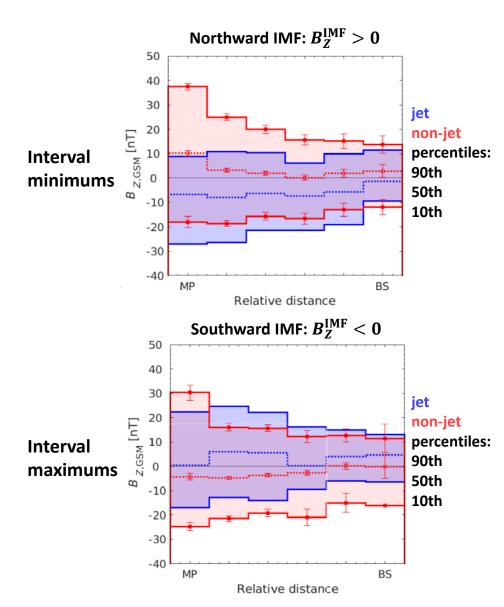




During northward IMF, more southward fields in jets close to the magnetopause

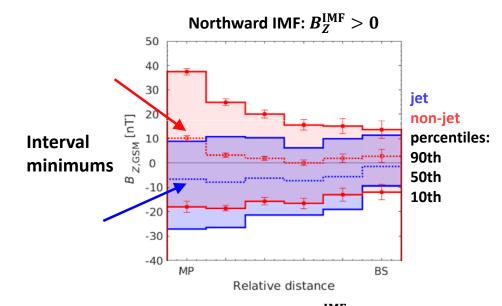
→ Favorable for enhancing reconnection

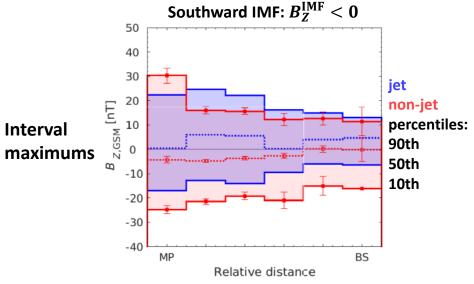
Are the variations in jets larger than in similar intervals within the ambient magnetosheath?



During northward IMF, jets exhibit more southward fields than similar magnetosheath intervals. Over half of the jets exhibit southward fields!

Are the variations in jets larger than in similar intervals within the ambient magnetosheath?





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Are the variations in jets larger than in similar intervals within the ambient magnetosheath?

During southward IMF, there is more variation within the similar magnetosheath intervals.

