

Velocity and E-field Variations of Geomagnetic Field During the Passage of Interplanetary Shocks

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- Previous works on velocity and electric field measurements
 - 1980s
 - s/c ISEE 1 and 2
 - No complete simultaneous measurements of different parameters like B, V
 - No high resolution measurements
 - No direct electric field measurements
 - E field is inferred from $E \times B$

Event Characterization

- IP identification

- Box type variation with a sharp rise in
 - Density
 - Velocity
 - Thermal Speed
 - Temperature
 - Dynamic pressure



- Box type variation
- Increase of about twice the background

- Themis A/D/E spacecraft

- Located within 45° of sun-earth line.
- Magnetic field
- Velocity and
- Electric field measurements

- GOES magnetic field

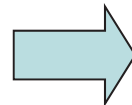
- Other spacecraft

- Cluster (?)
- Geotail

- Ground magnetometers

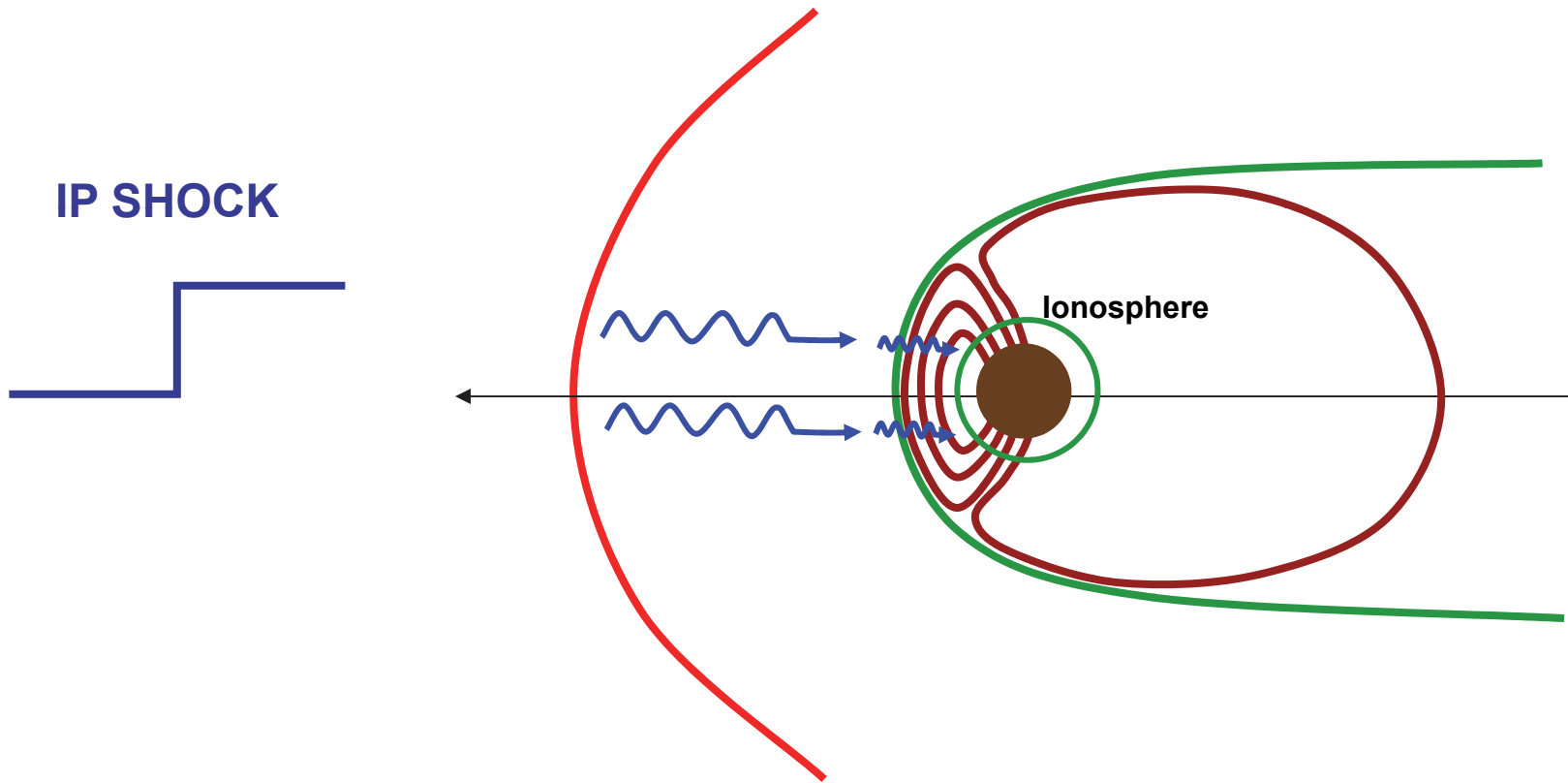
IP shock interaction with magnetosphere

- Interaction with bow shock
- Transmission in the magnetosheath
- Interaction with magnetopause
- Transmission of IP shock into the magnetosphere
 - Mpause inward motion in response to IP shock in the sheath drives a fast wave into the magnetosphere
- **Reflected wave concept ??**
 - **MHD models**
 - **LANL observations**



- Modifications of field aligned and ionospheric current systems
- Magnetic disturbances on the ground

IP shock and magnetosphere



Procedure

- Find IP shock

→ study what Themis sees

magnetopause motion

magnetic field perturbations

velocity and E-field variations

ground signatures

etc.

IP shock



Magnetospheric
consequence

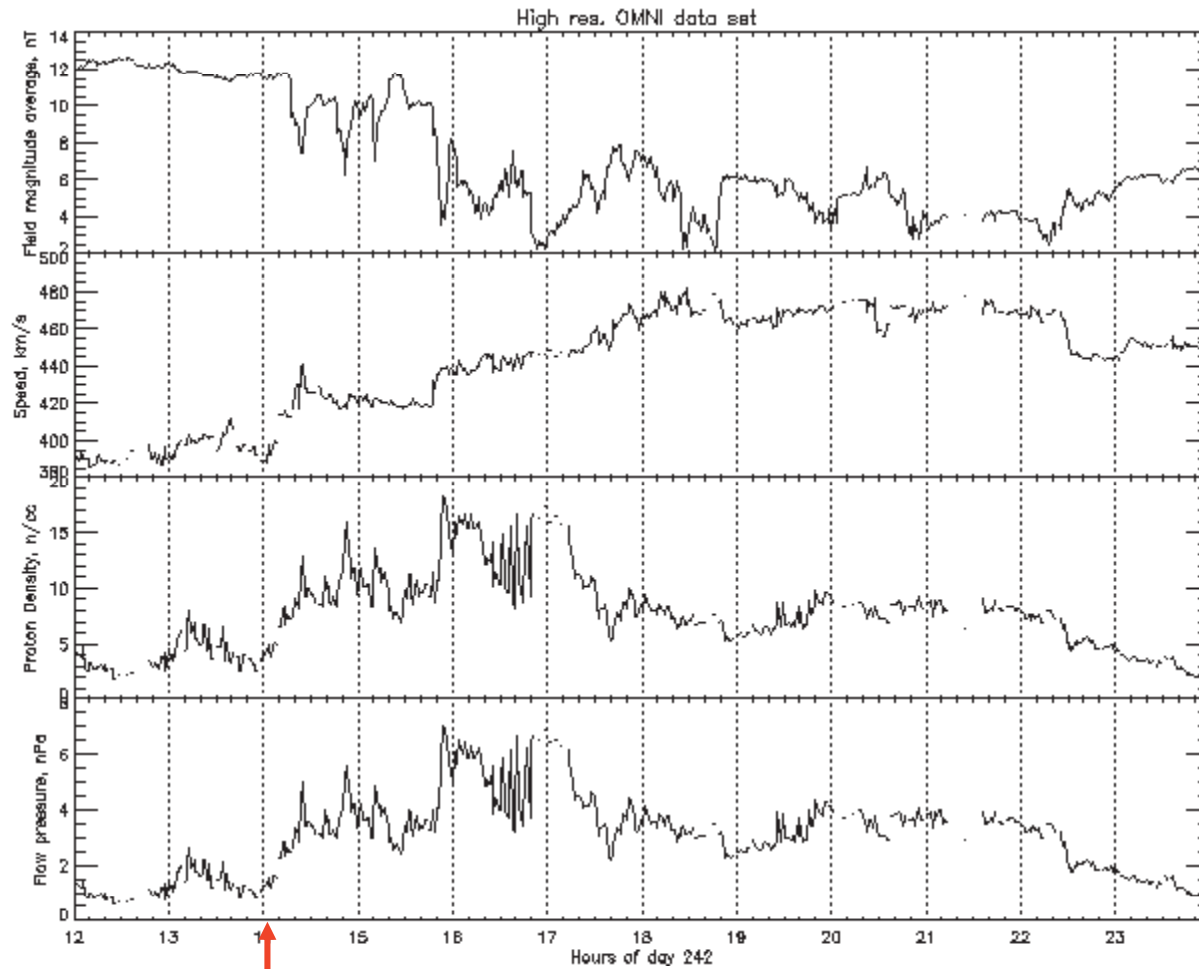
Case-1: August 30, 2009

Btot

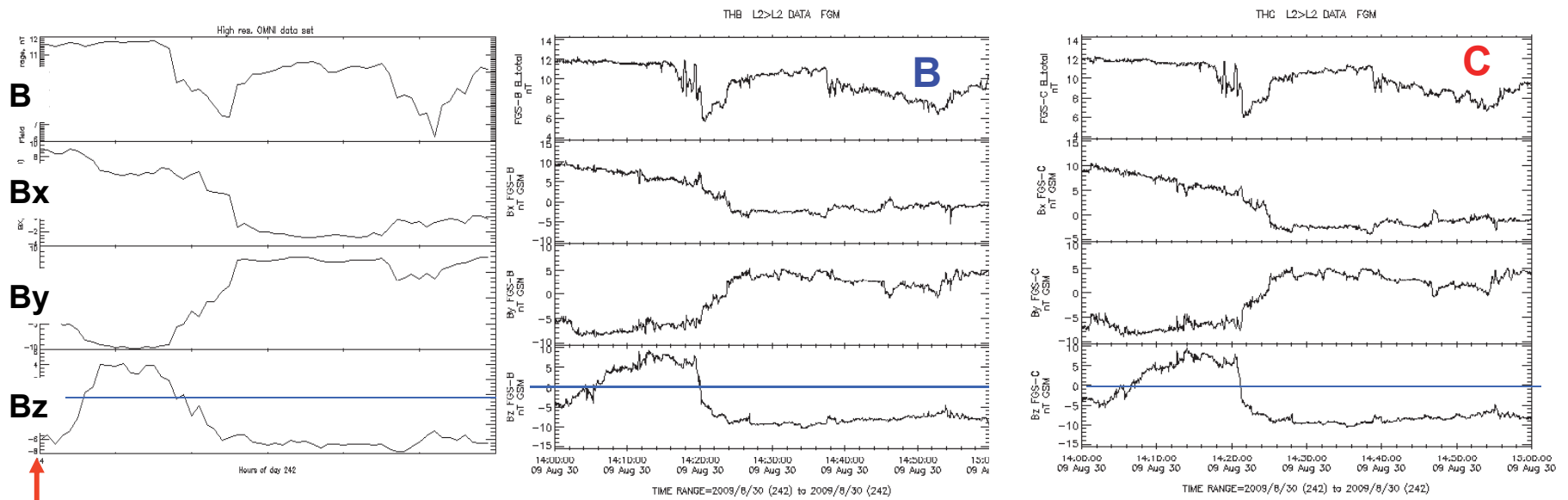
Vsw

Nsw

Pdyn

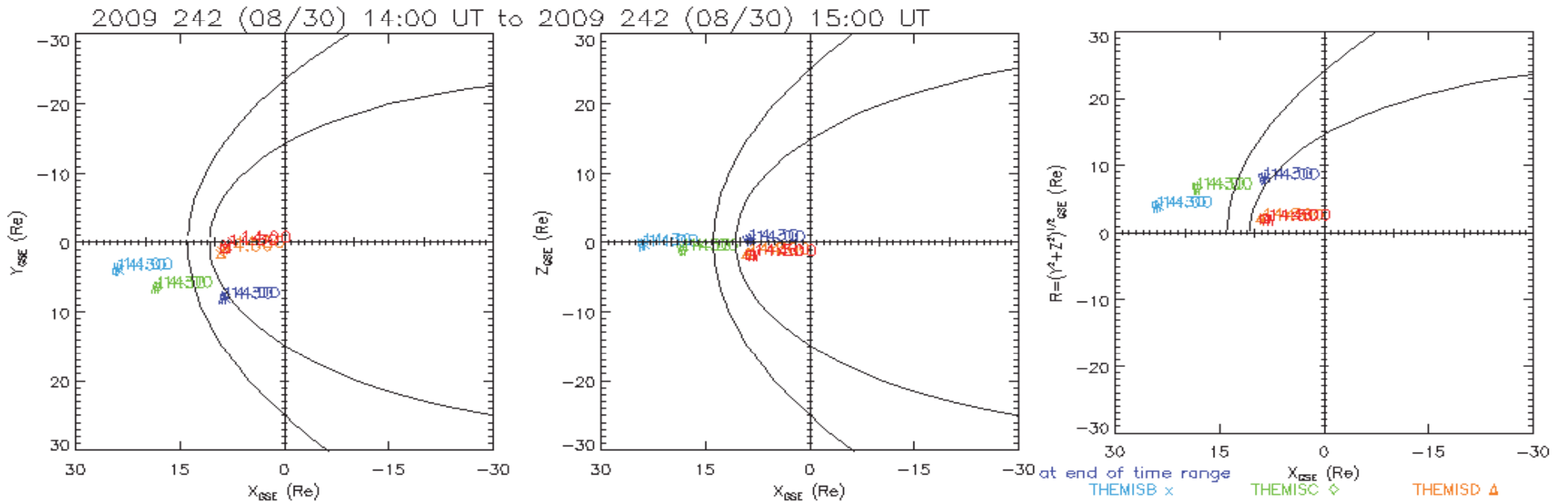


Interplanetary Magnetic Field



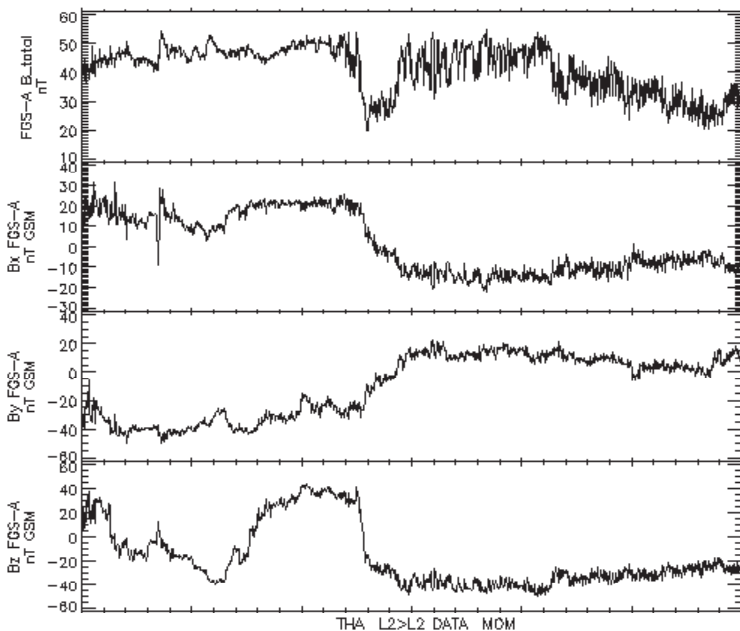
Please acknowledge data provider, V. Angelopoulos, U. Auster & K.H. Glasmeier and W. Bai at UCB, TUBS and IWF respectively, NASA NASS-02099 and CDAWeb when using these data. Generated by CDAWeb on Sun Sep 11 14:24:17 2011

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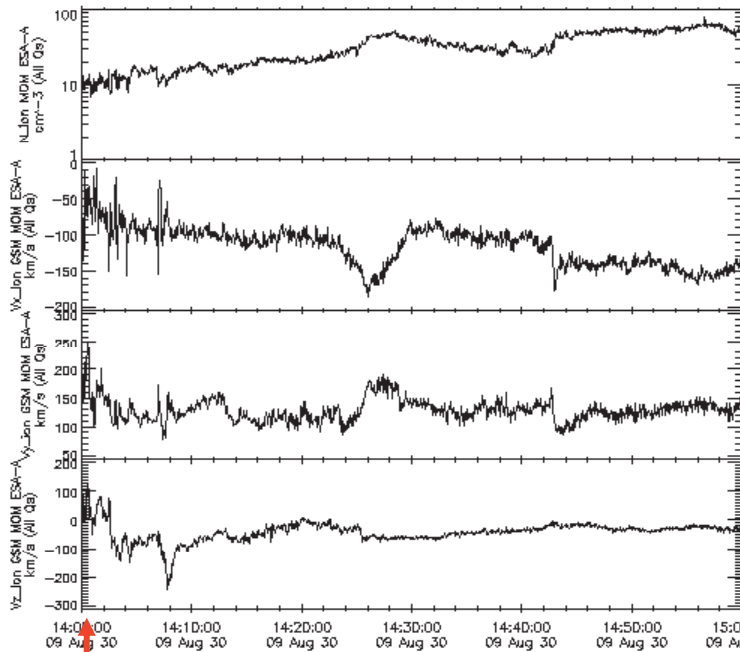


TA

THA L2>L2 DATA FGM



THA L2>L2 DATA MOM

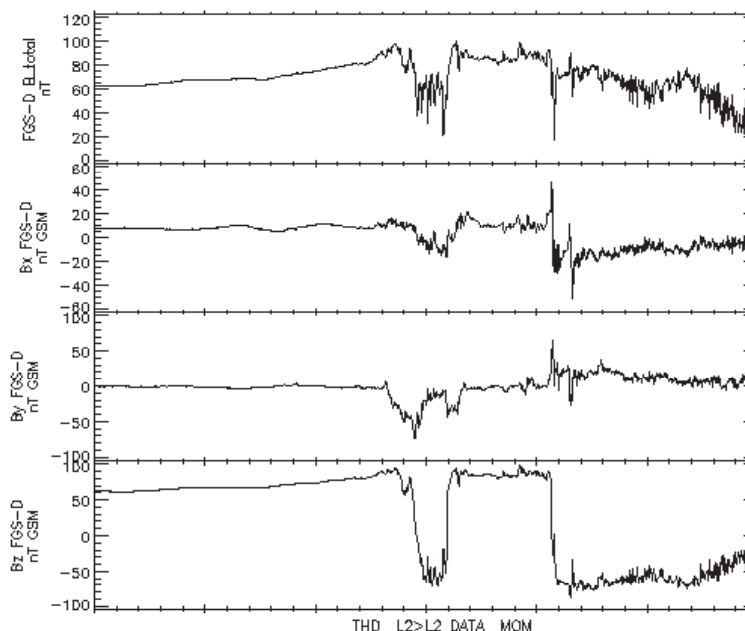


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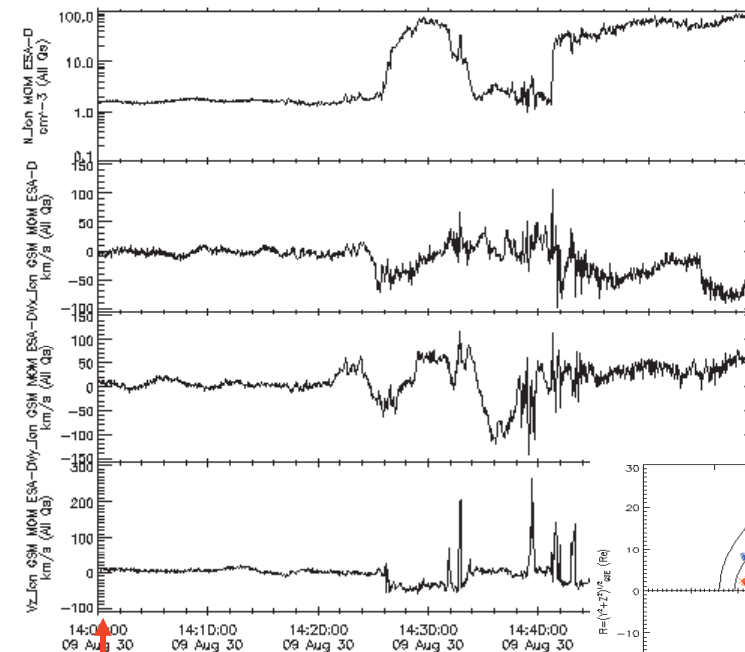
Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NAS5-02099 and CDAWeb when using these data.
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TD

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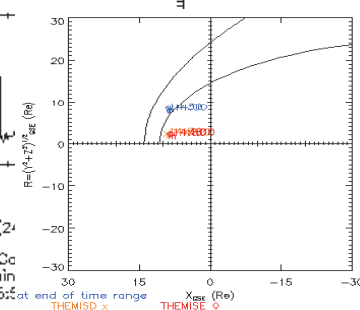


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TIME RANGE=2009/8/30 (242) to 2009/8/30 (242)

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TD

TE

Bx

By

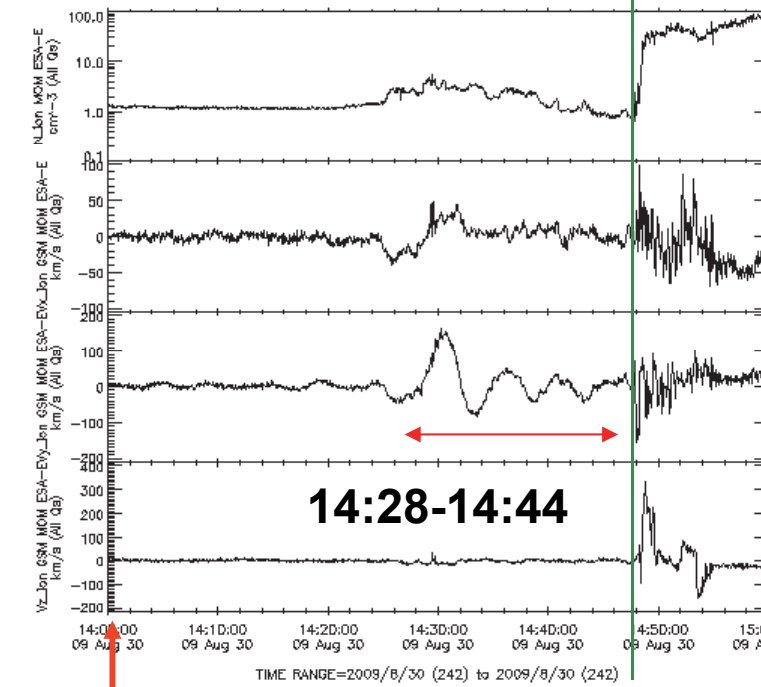
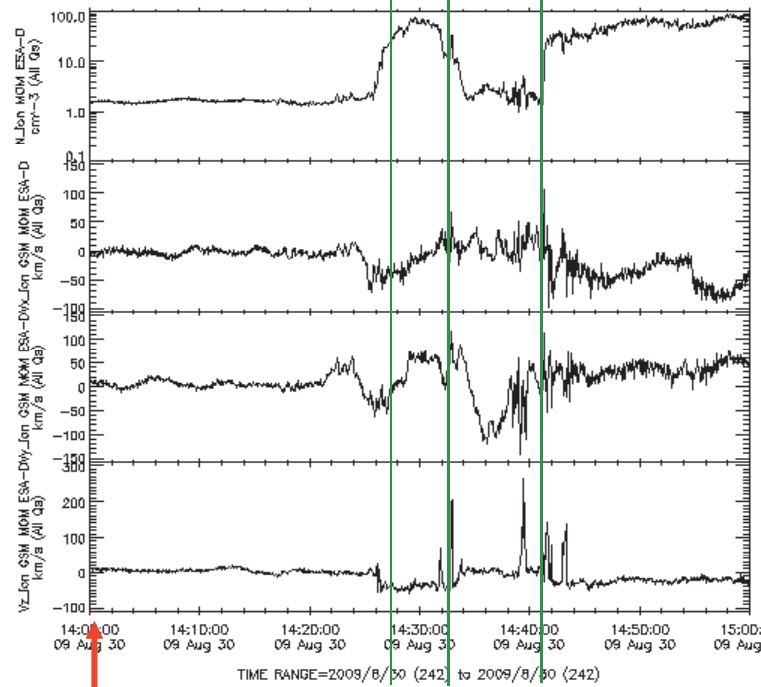
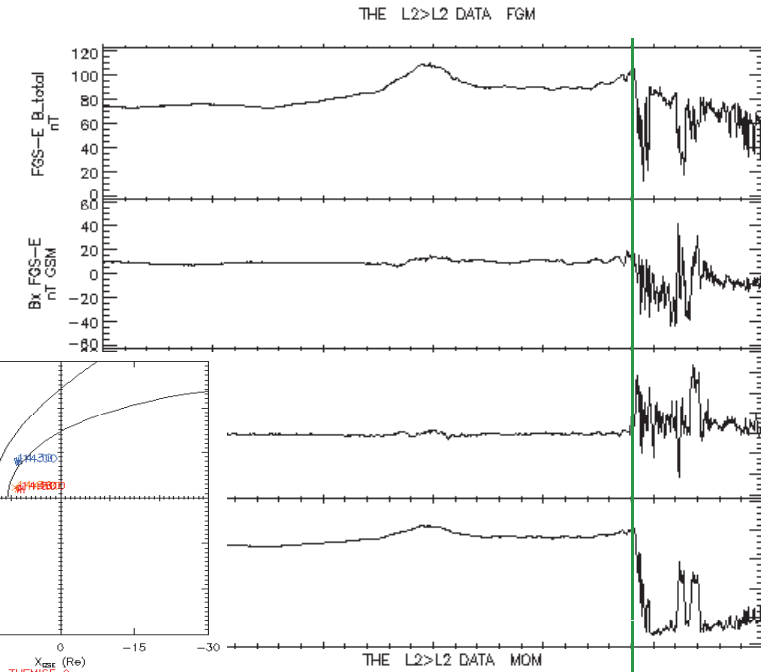
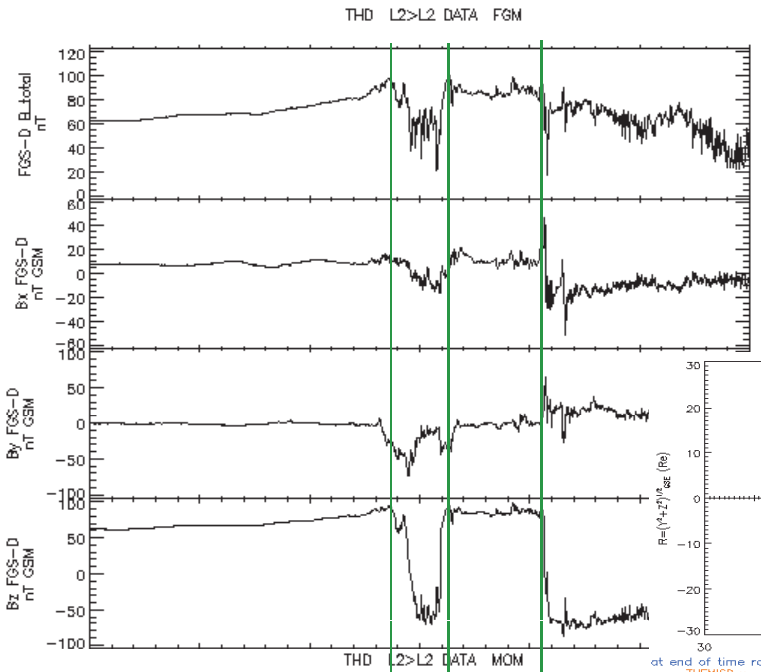
Bz

n

Vx

Vy

Vz



Vy

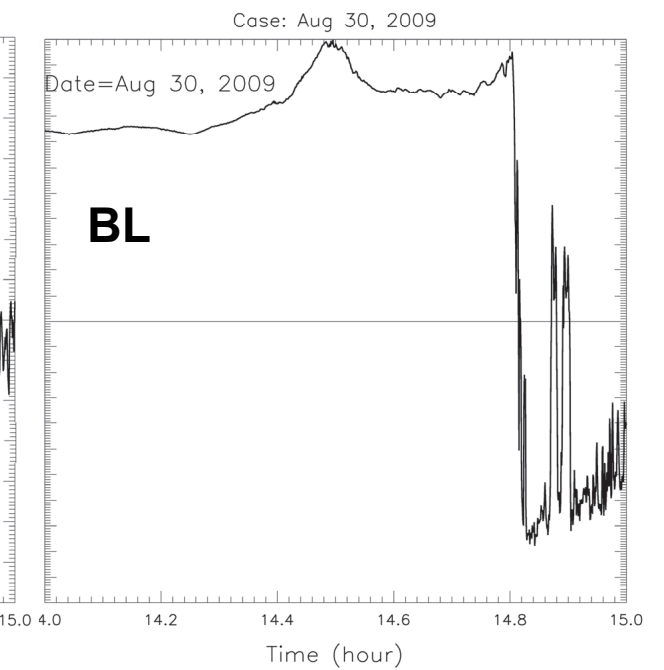
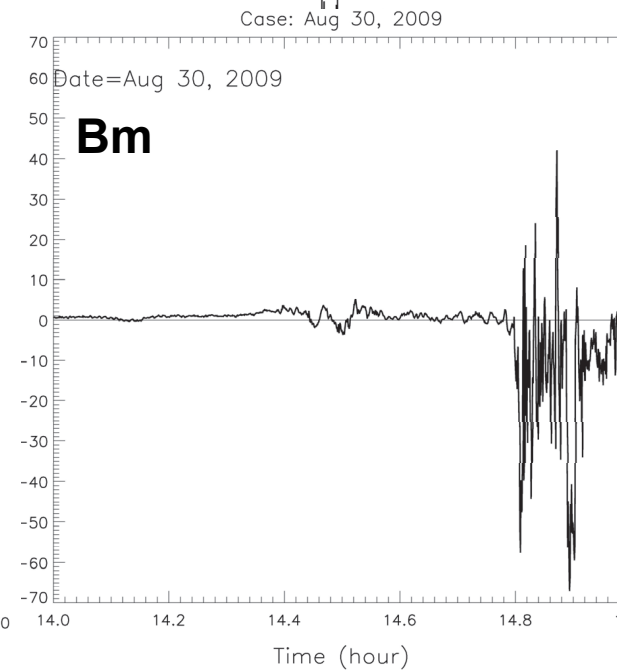
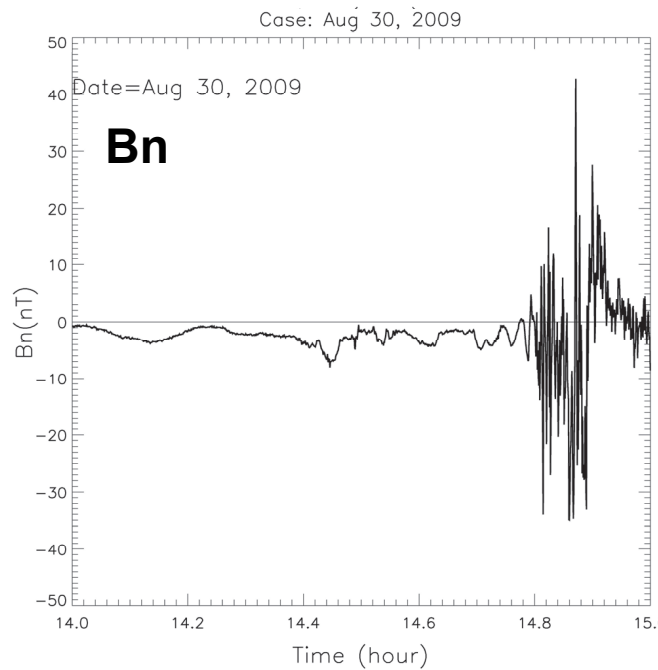
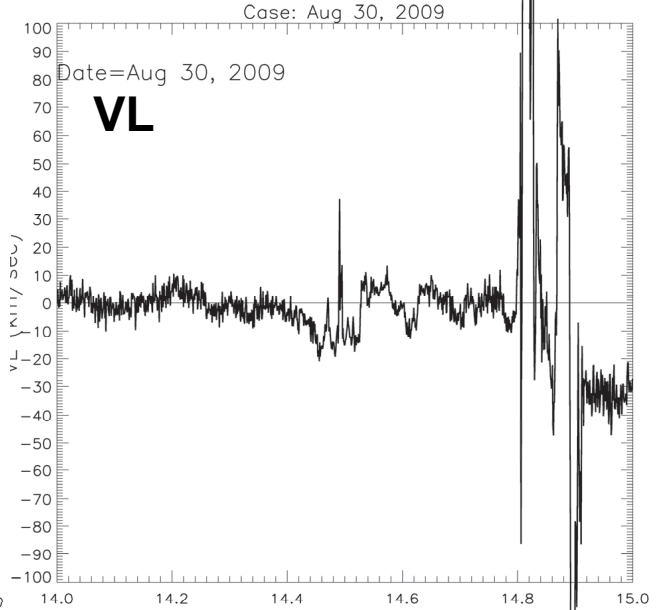
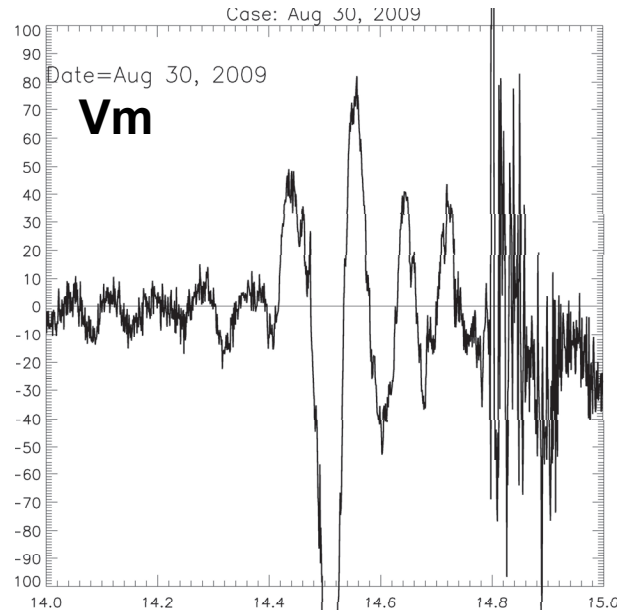
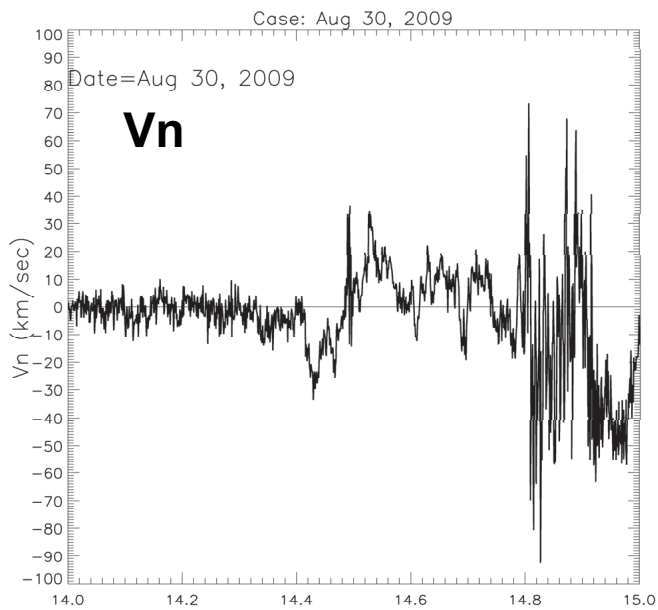
16min

14:28-14:44

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LMN-boundary coordinates: THE



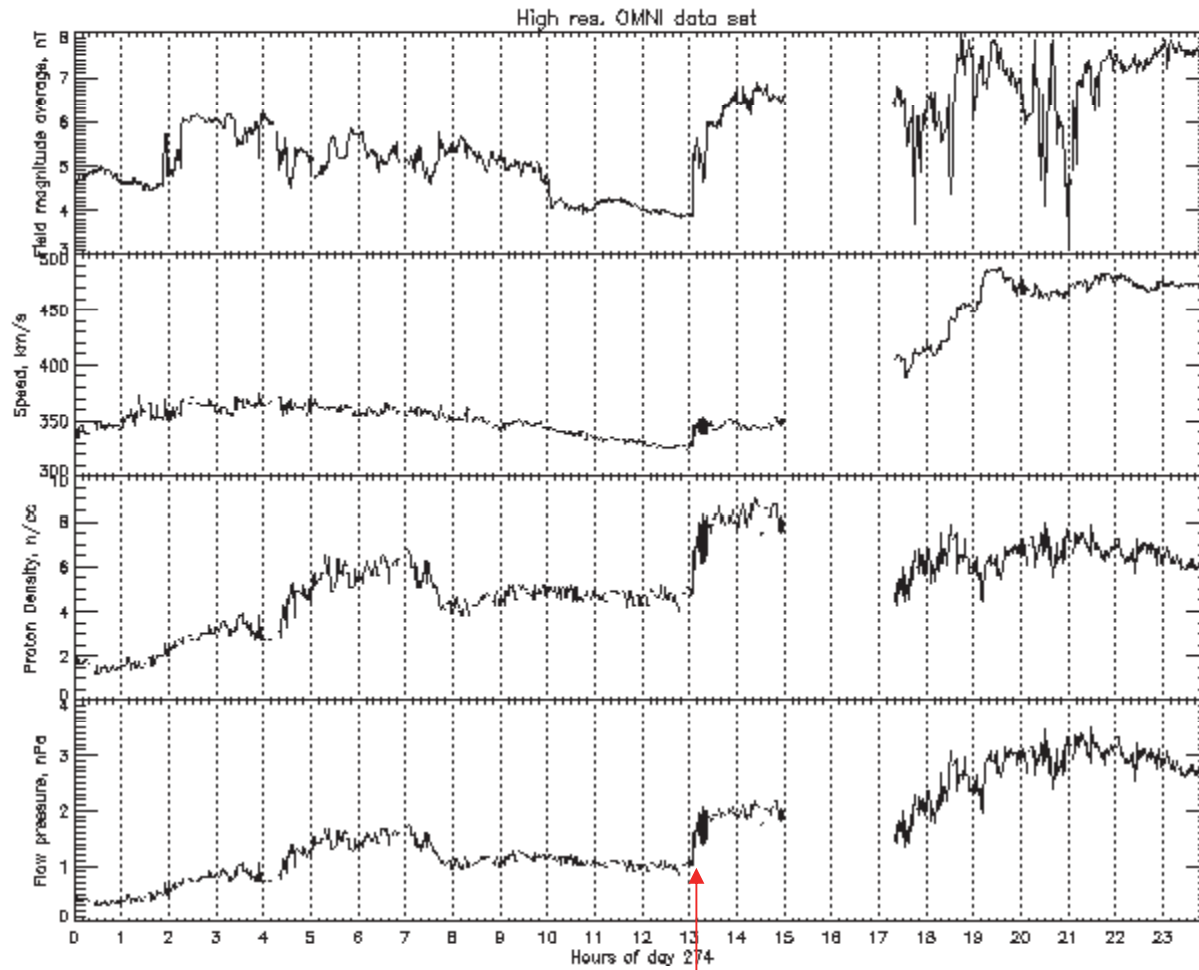
Case 2: Sep 30, 2008

Btot

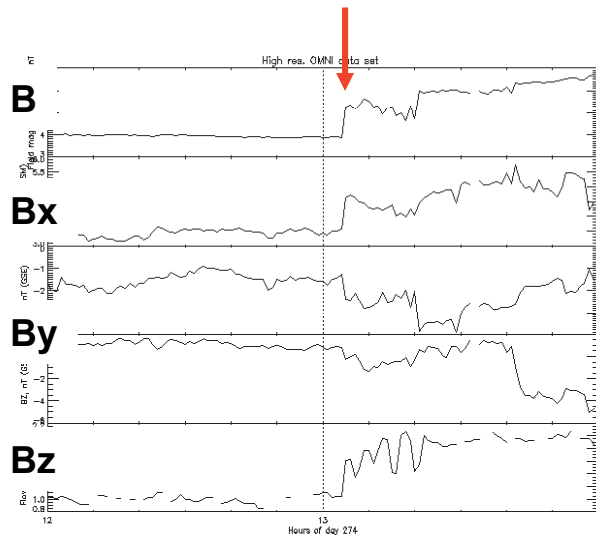
Vsw

Nsw

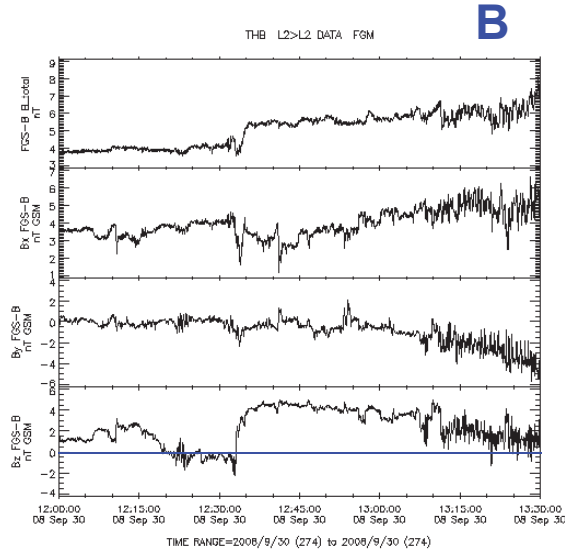
Pdyn



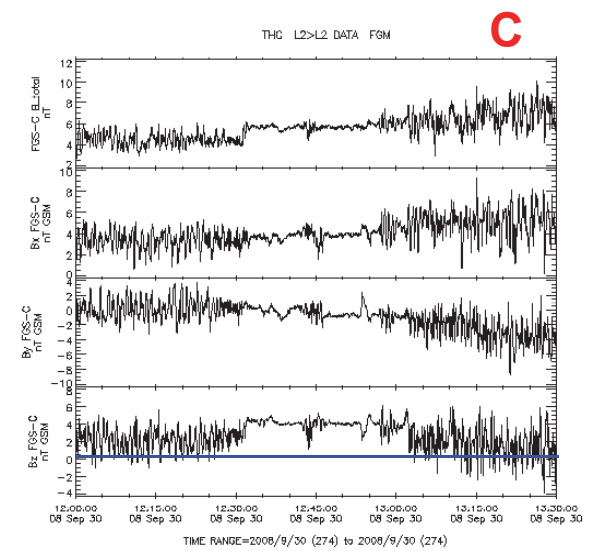
Interplanetary Magnetic Field



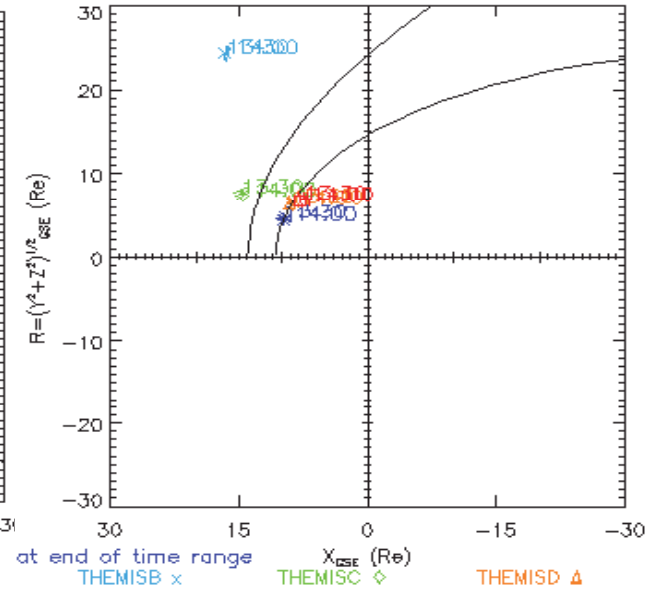
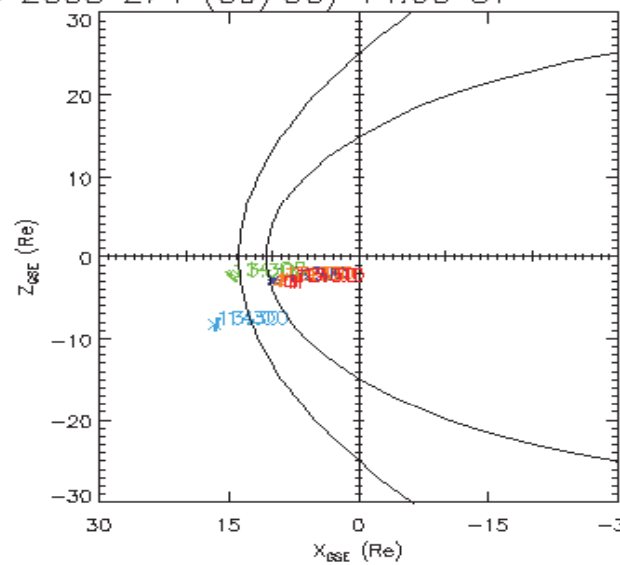
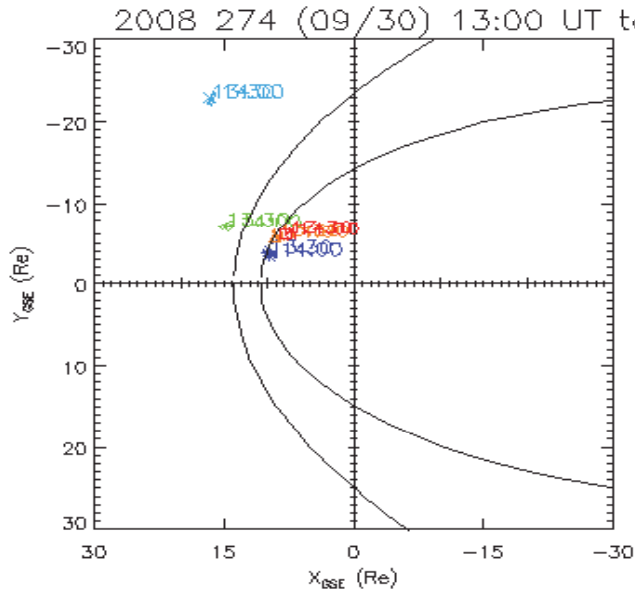
Northward IMF Bz



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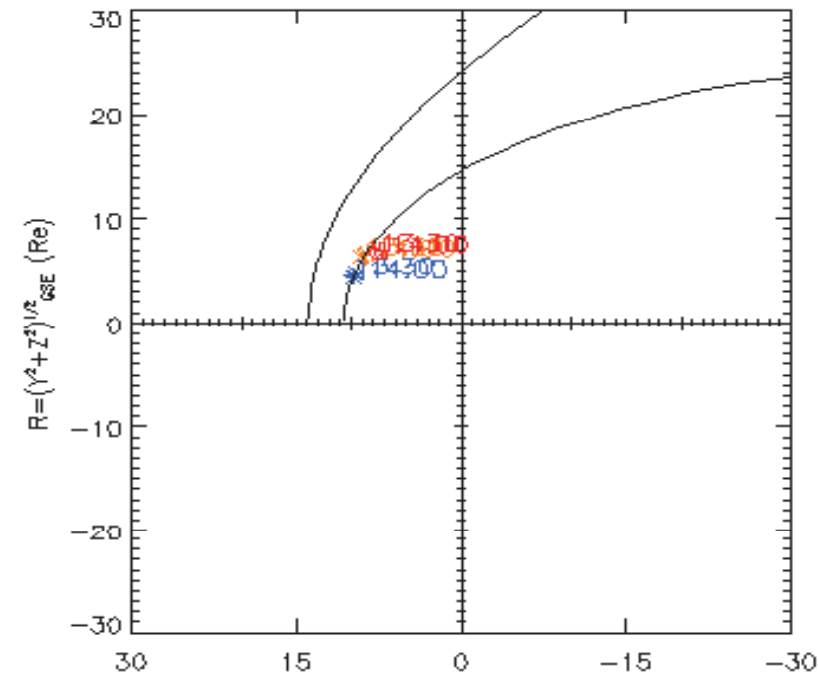
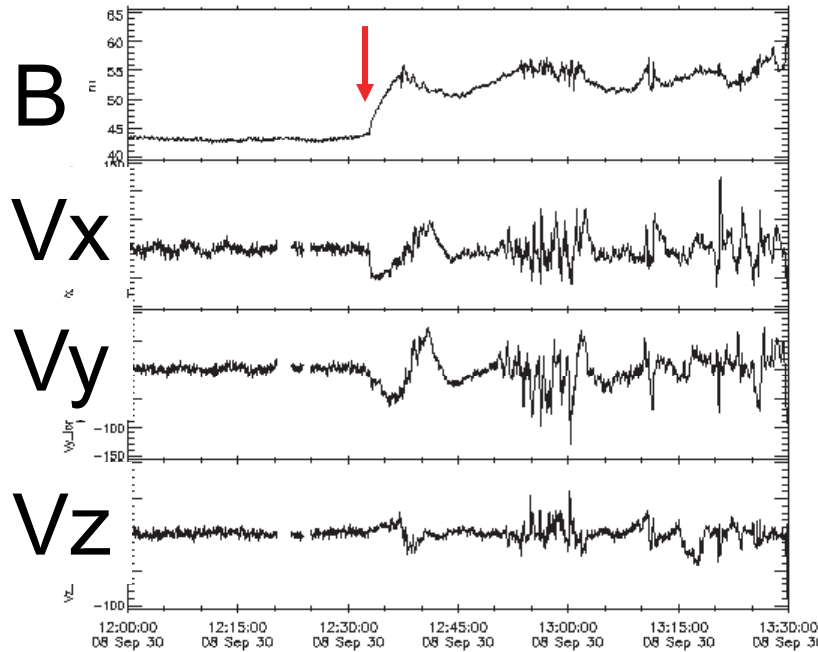


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TA

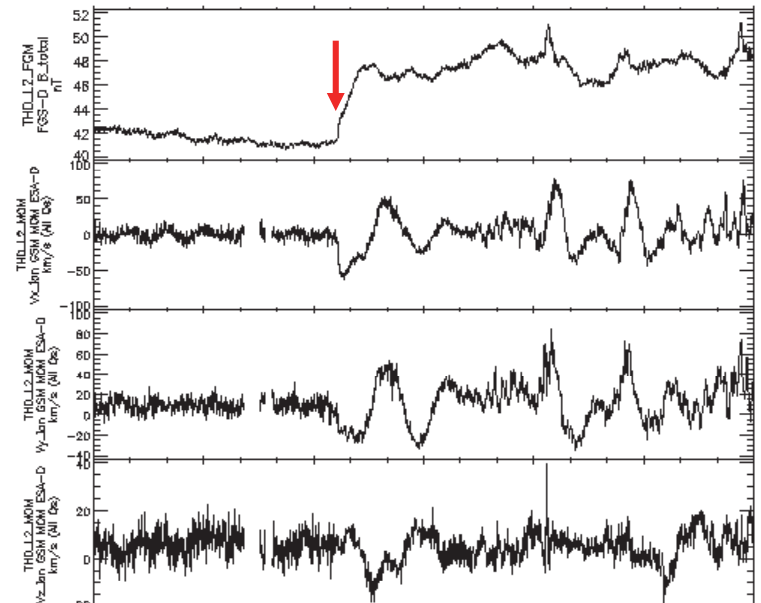
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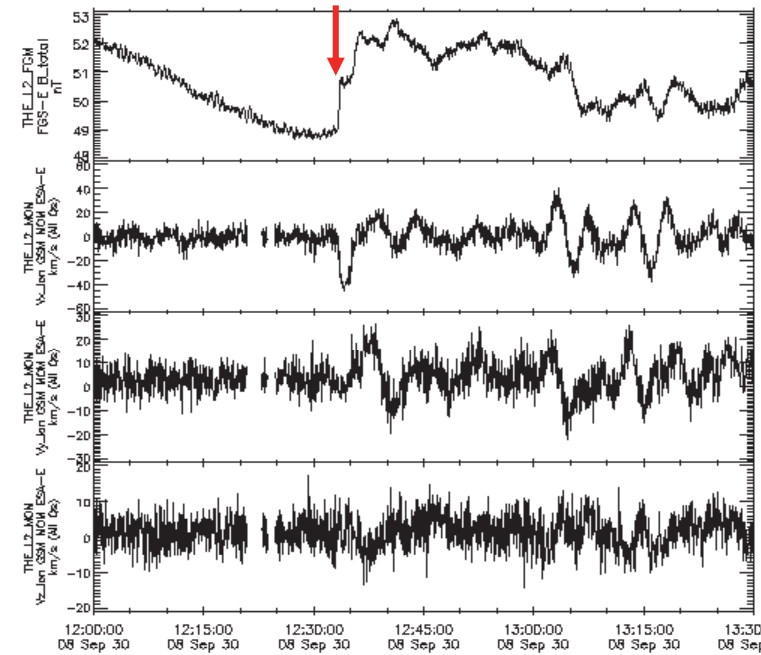
at end of time range
 THEMISD × THEMISE ◇

TD

Multiple datasets being plotted; refer to labels on either side of plot.



Multiple datasets being plotted; refer to labels on either side of plot.



TE

TIME RANGE=2008/9/30 (274) to 2008/9/30 (274)
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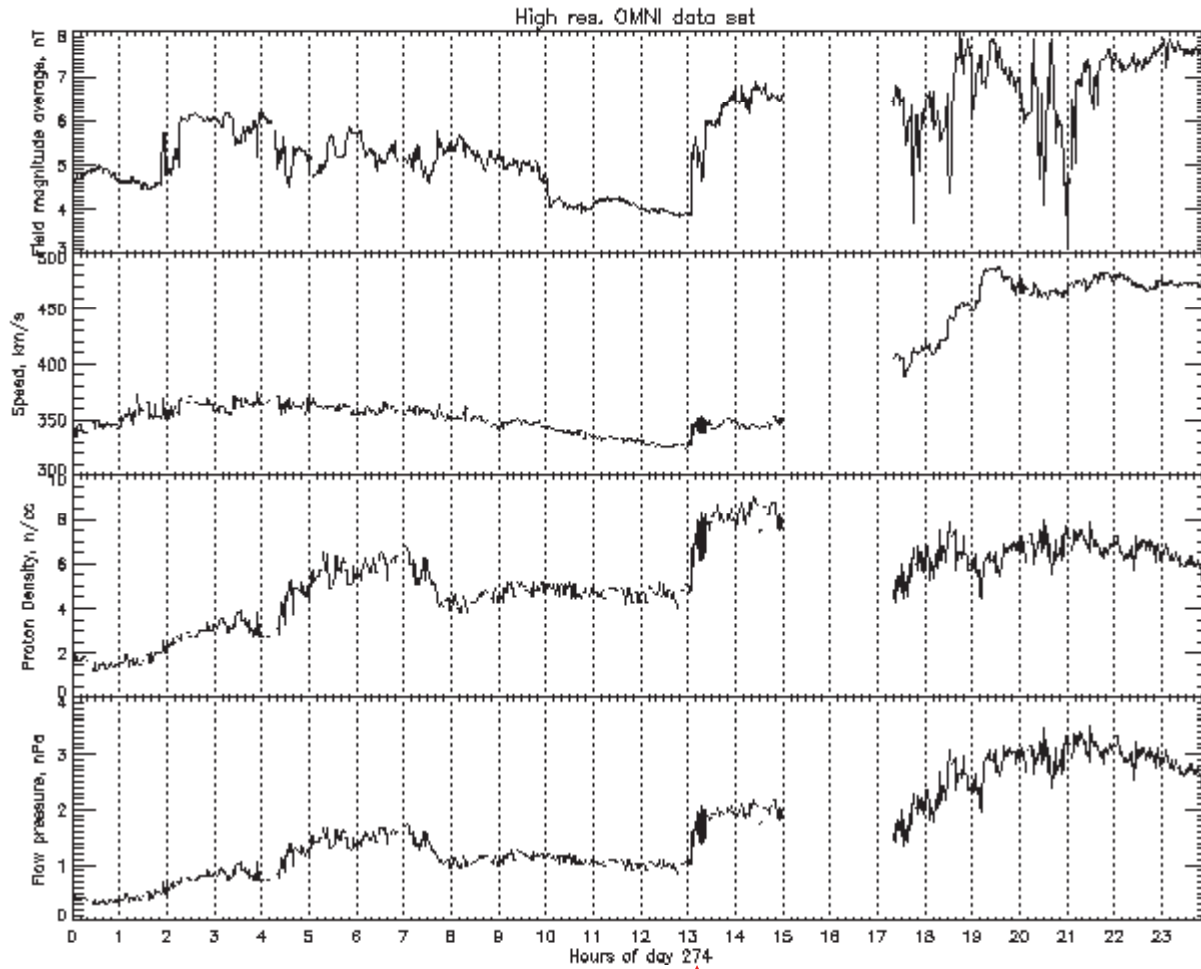
Case 3: July 22, 2008

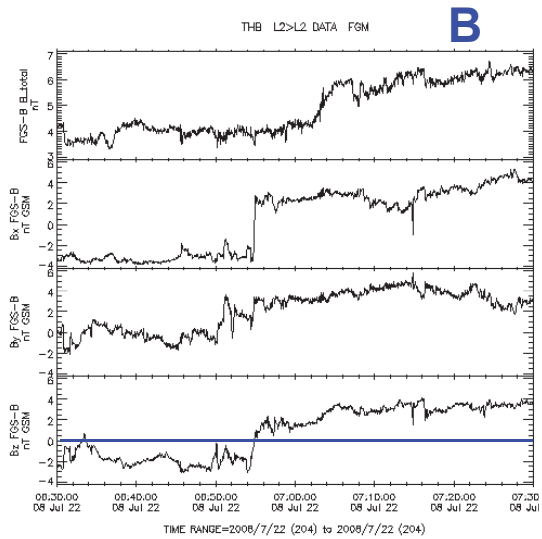
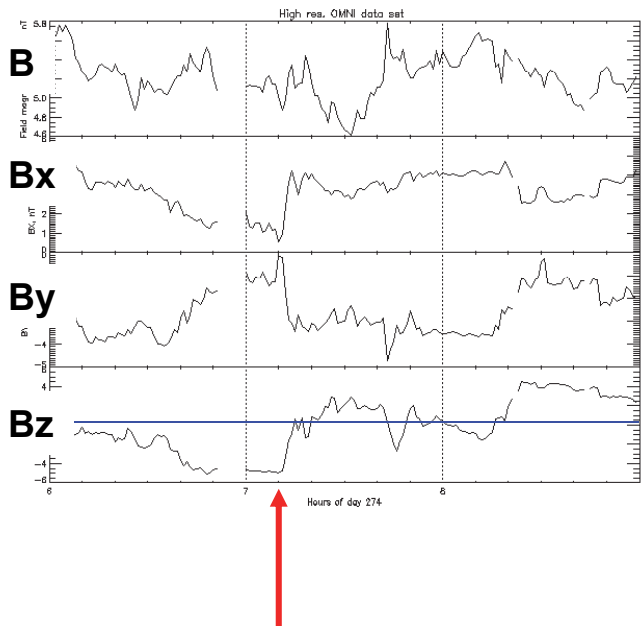
Btot

Vsw

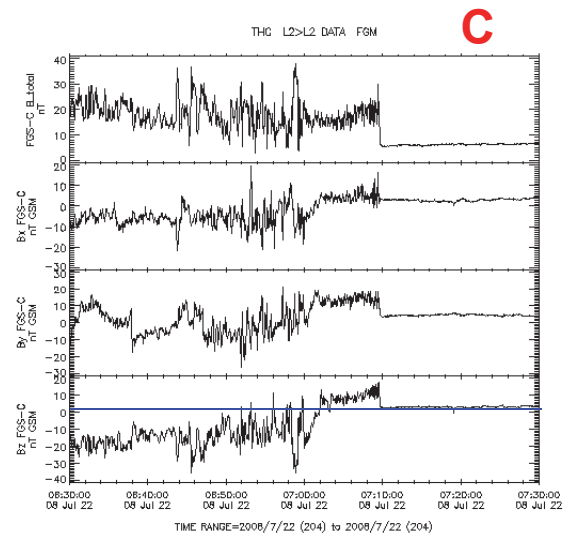
Nsw

Pdyn

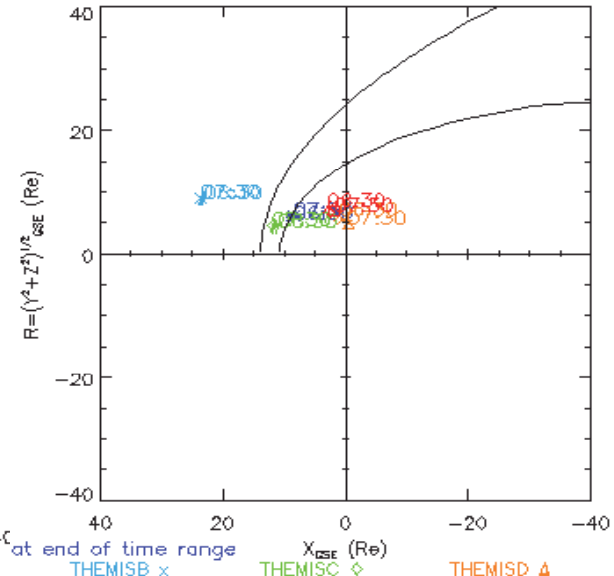
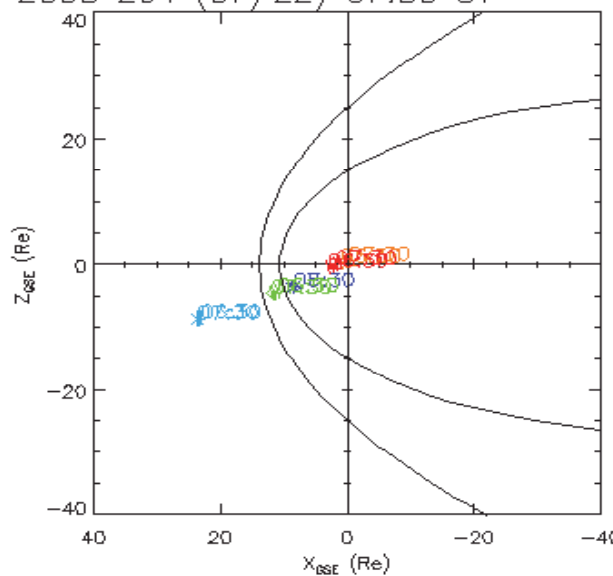
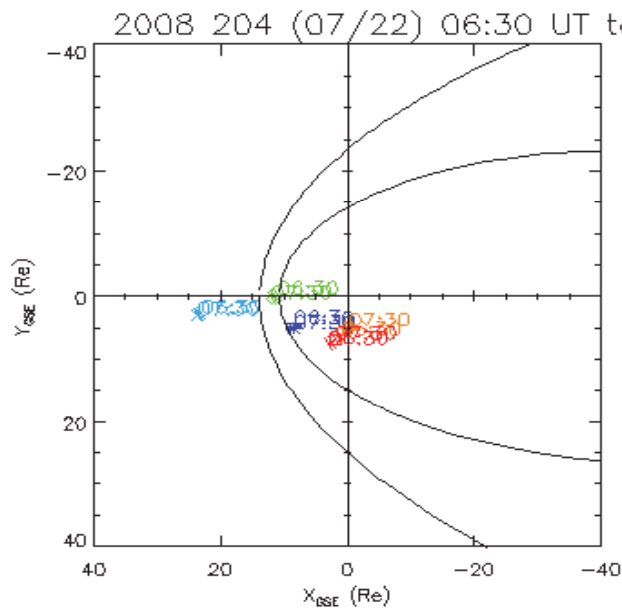




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TA

Bx

By

Bz

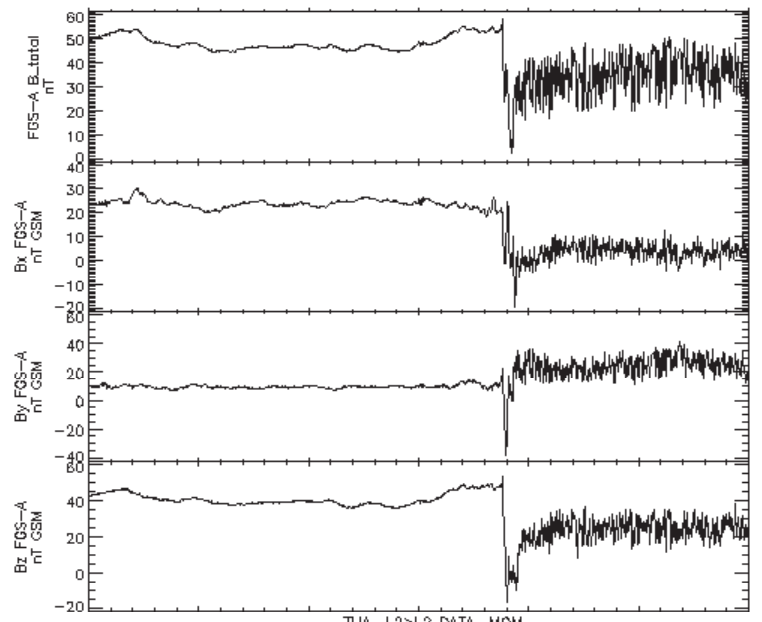
n

Vx

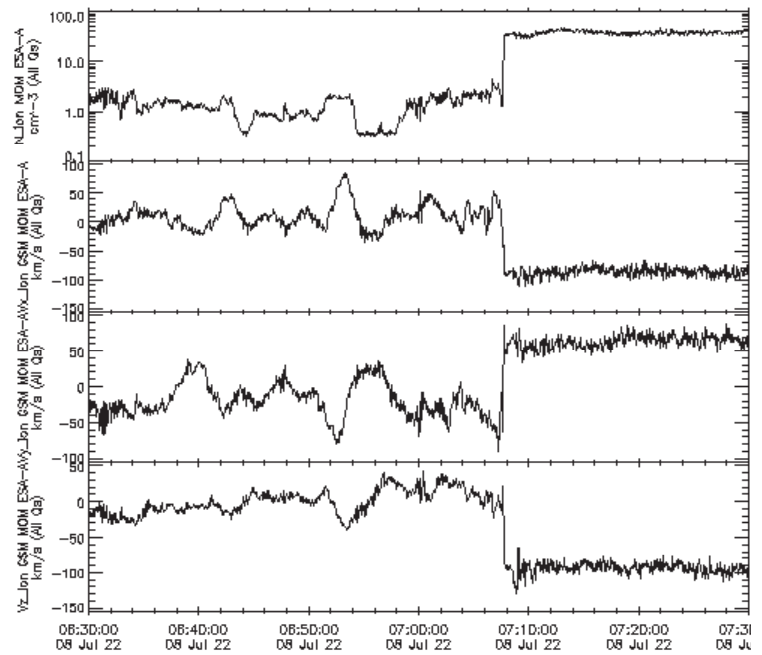
Vy

Vz

THA L2>L2 DATA FGM



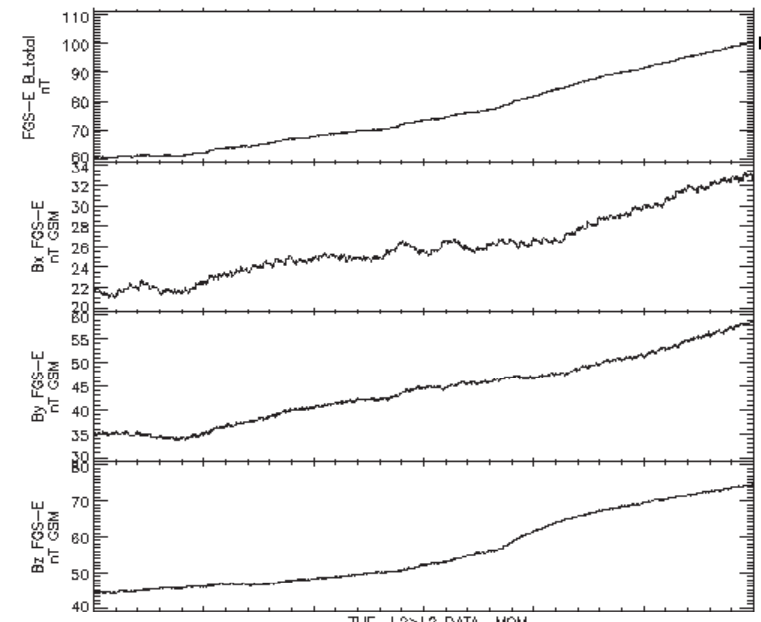
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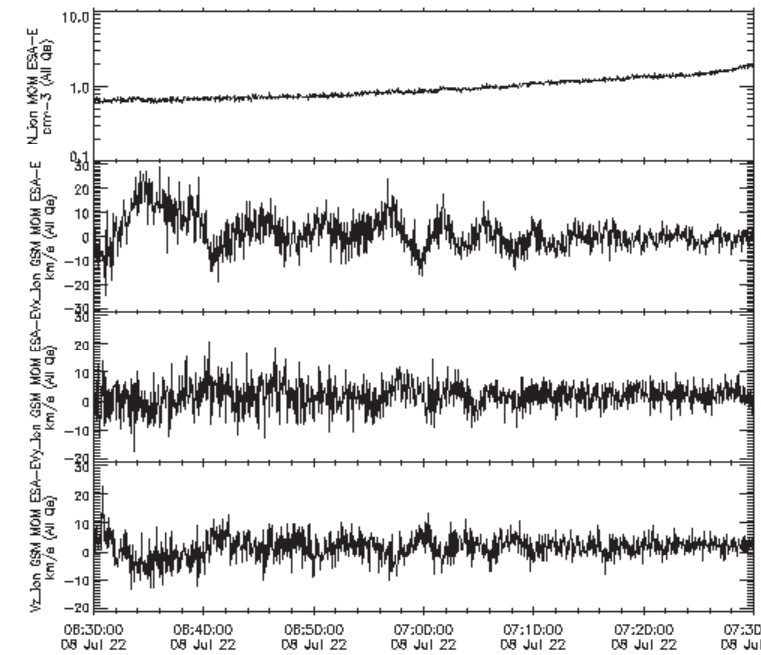
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THE L2>L2 DATA FGM



THE L2>L2 DATA MOM



TIME RANGE=2008/7/22 (204) to 2008/7/22 (204)

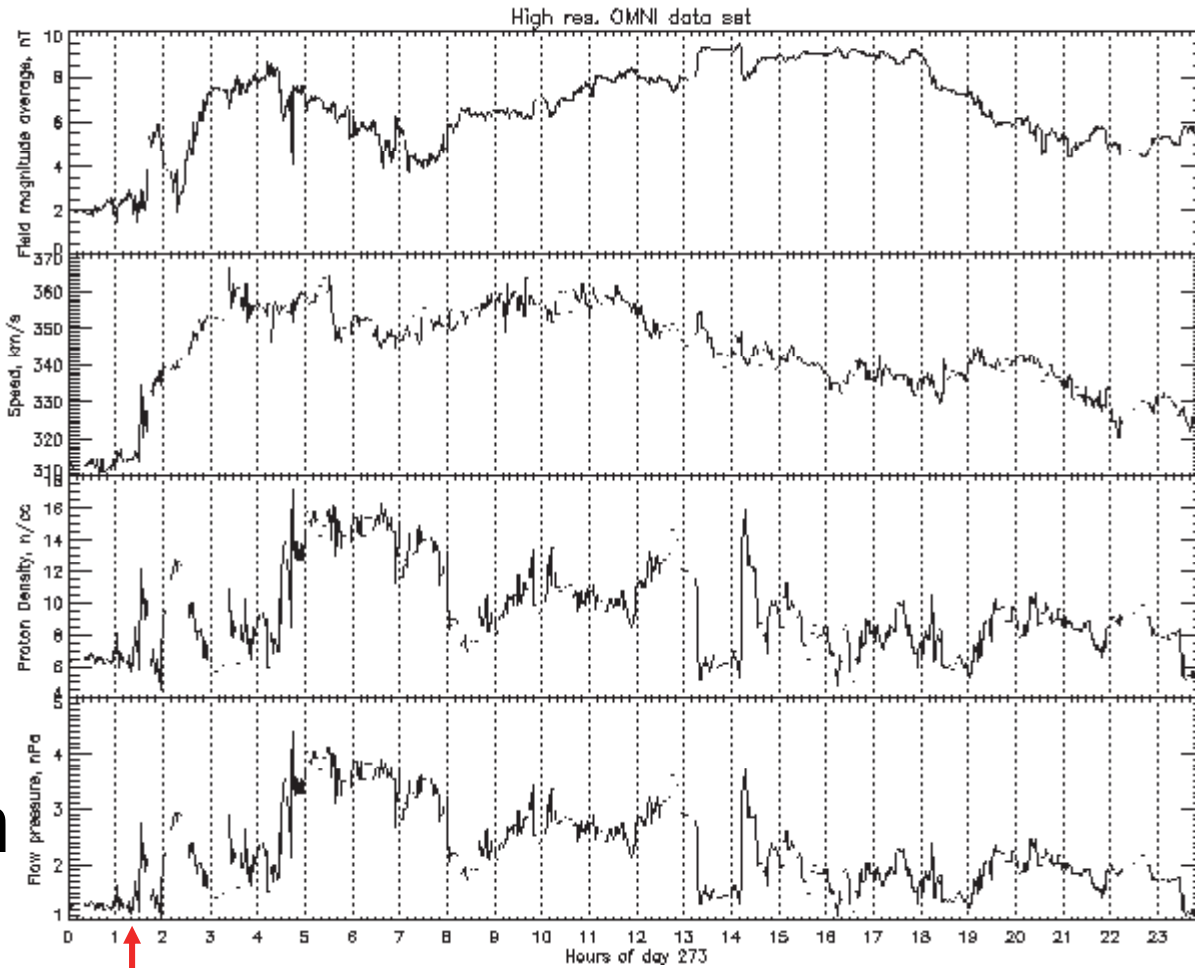
Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDWeb when using these data.
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TE

- Case 4: multiple shocks
- Case 5: s/c goes into sheath
- Case 6: A does not see anything but D and E sees periodic variations

Case 4: Sep 30, 2009

Btot



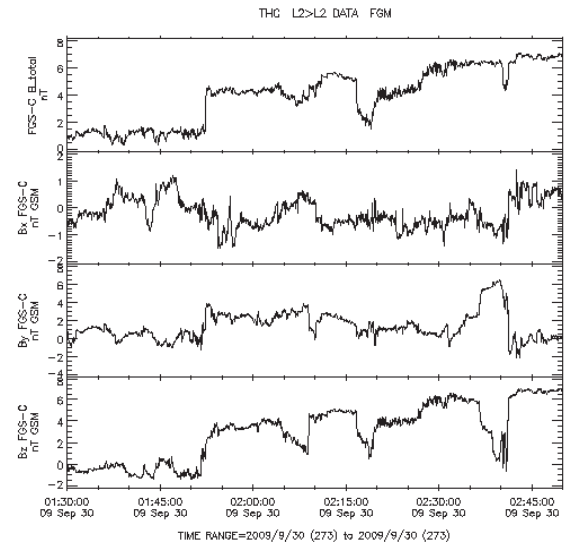
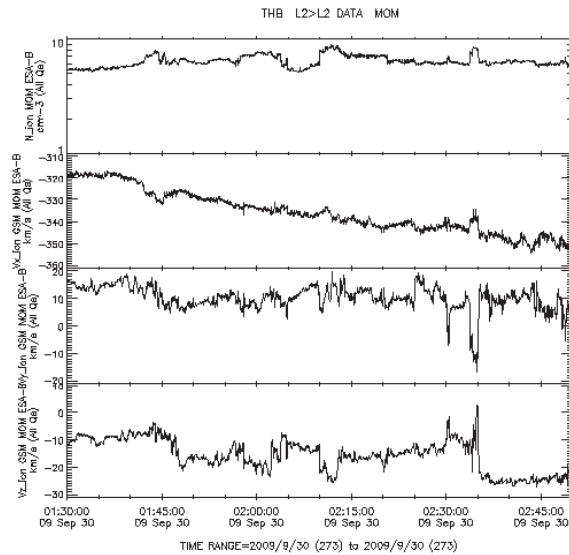
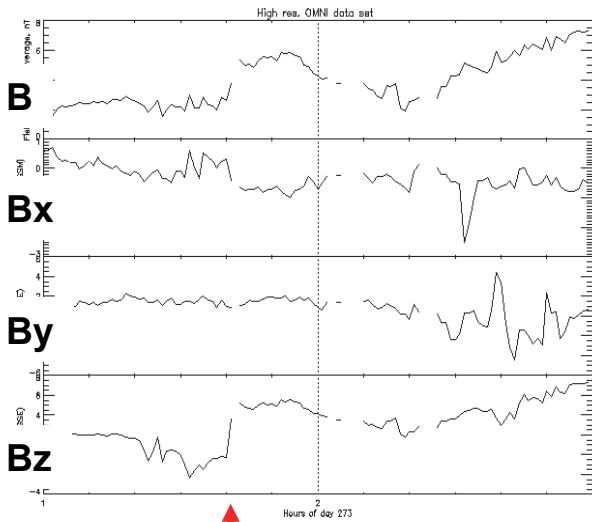
Vsw

Nsw

Pdyn

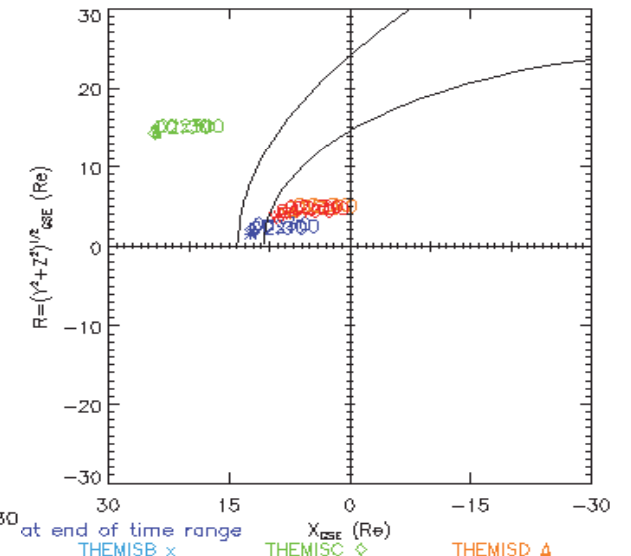
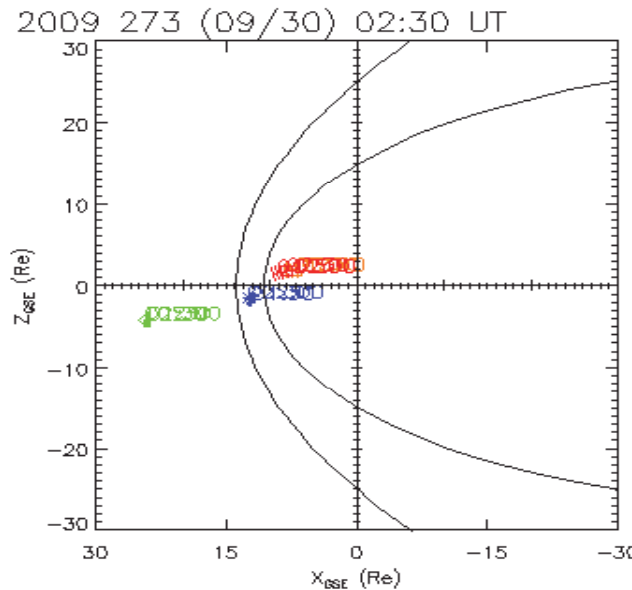
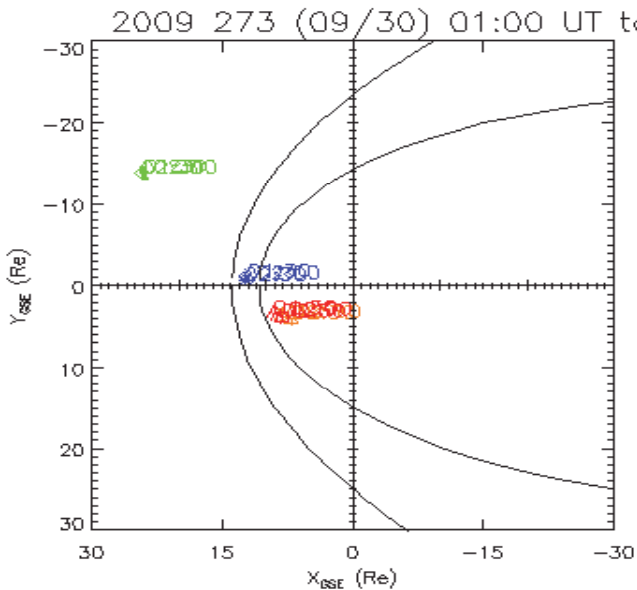
B, mom

C fgm



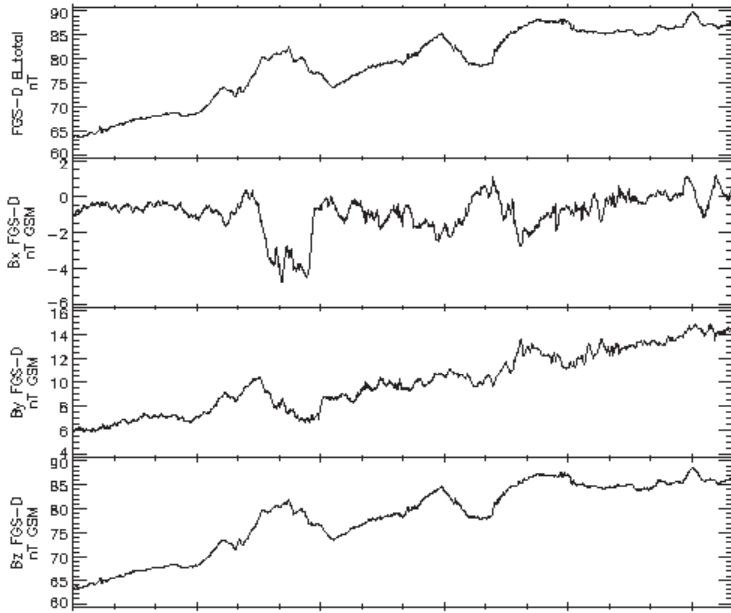
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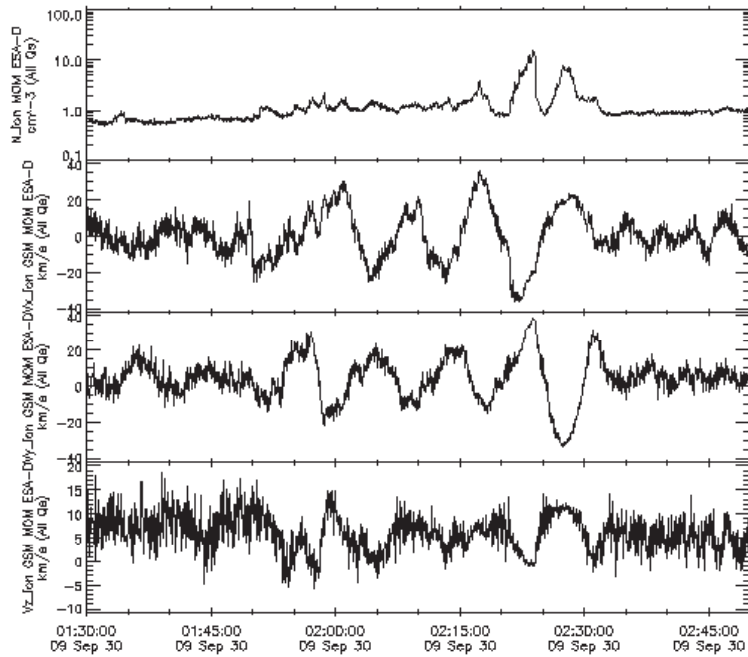


TD

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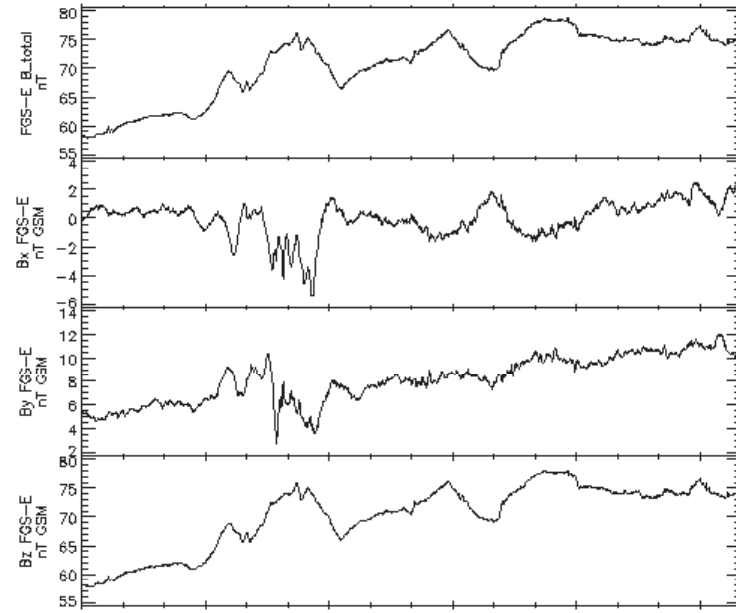
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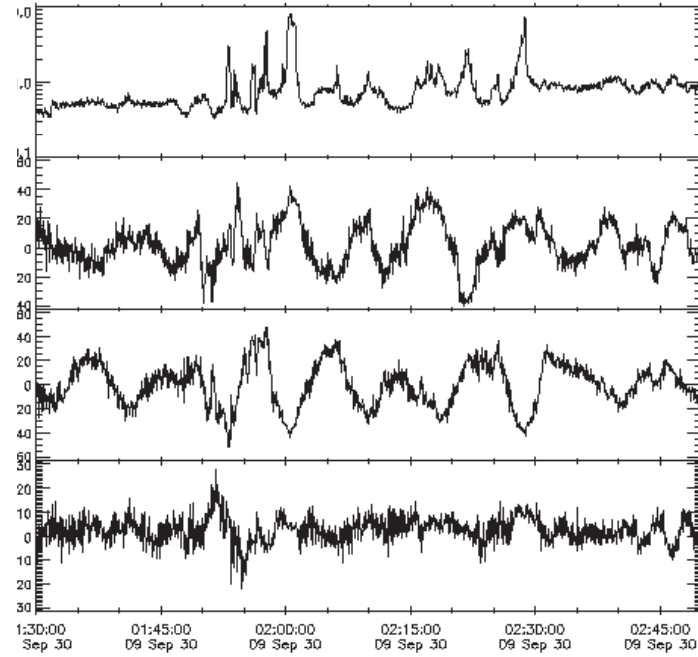
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Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDAWeb when using these data.
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THE L2>L2 DATA FGM



THE L2>L2 DATA MOM

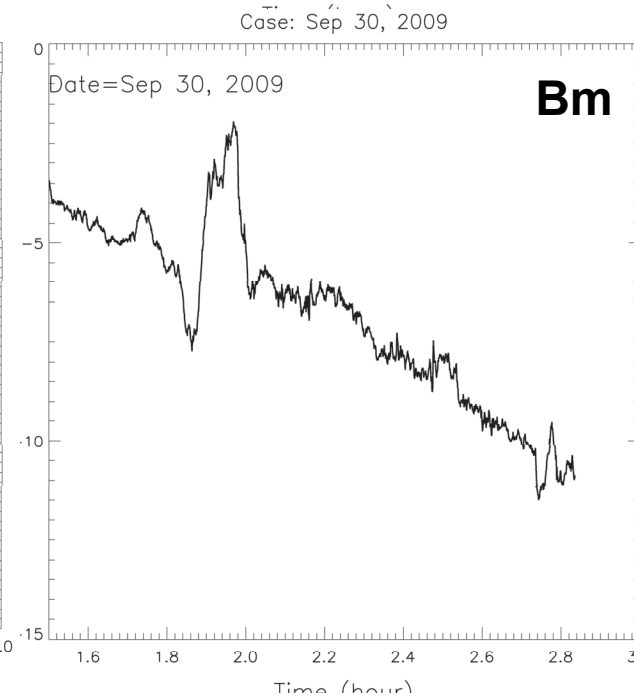
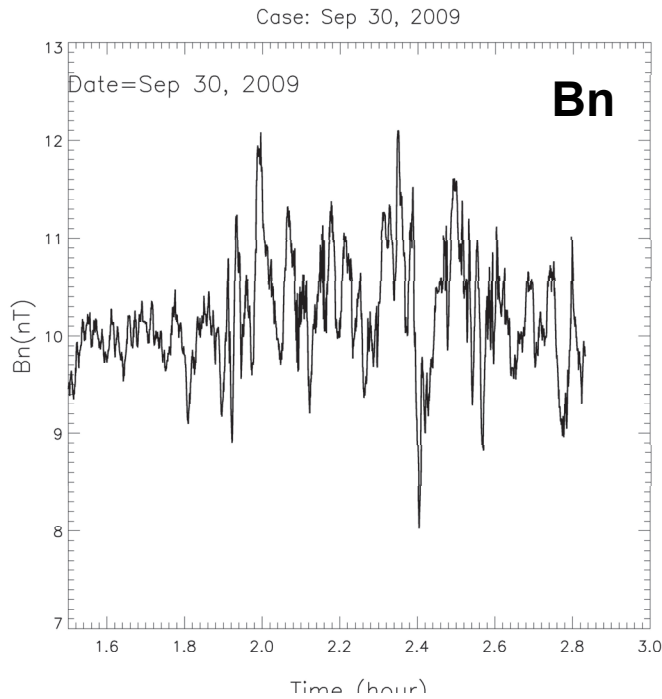
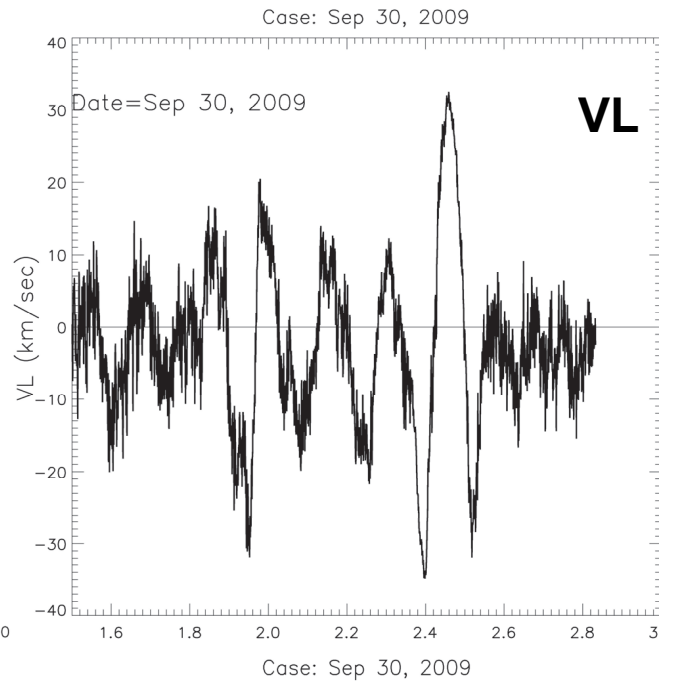
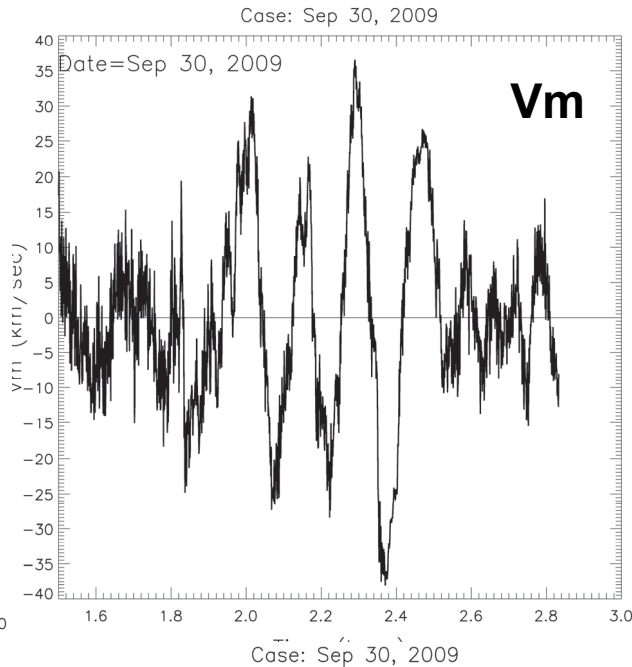
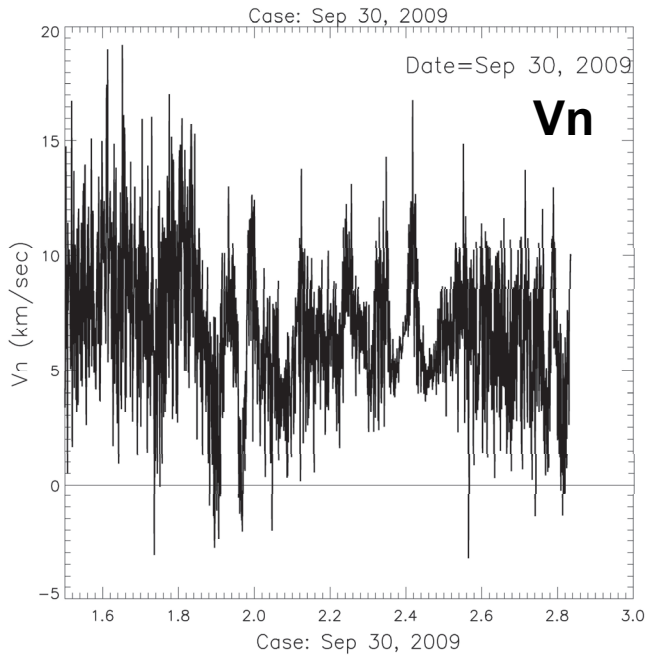


TIME RANGE=2009/9/30 (273) to 2009/9/30 (273)

Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDAWeb when using these data.
Generated by CDAWeb on Mon Sep 12 10:15:46 2011

TE

LMN Boundary Coordinates: THD



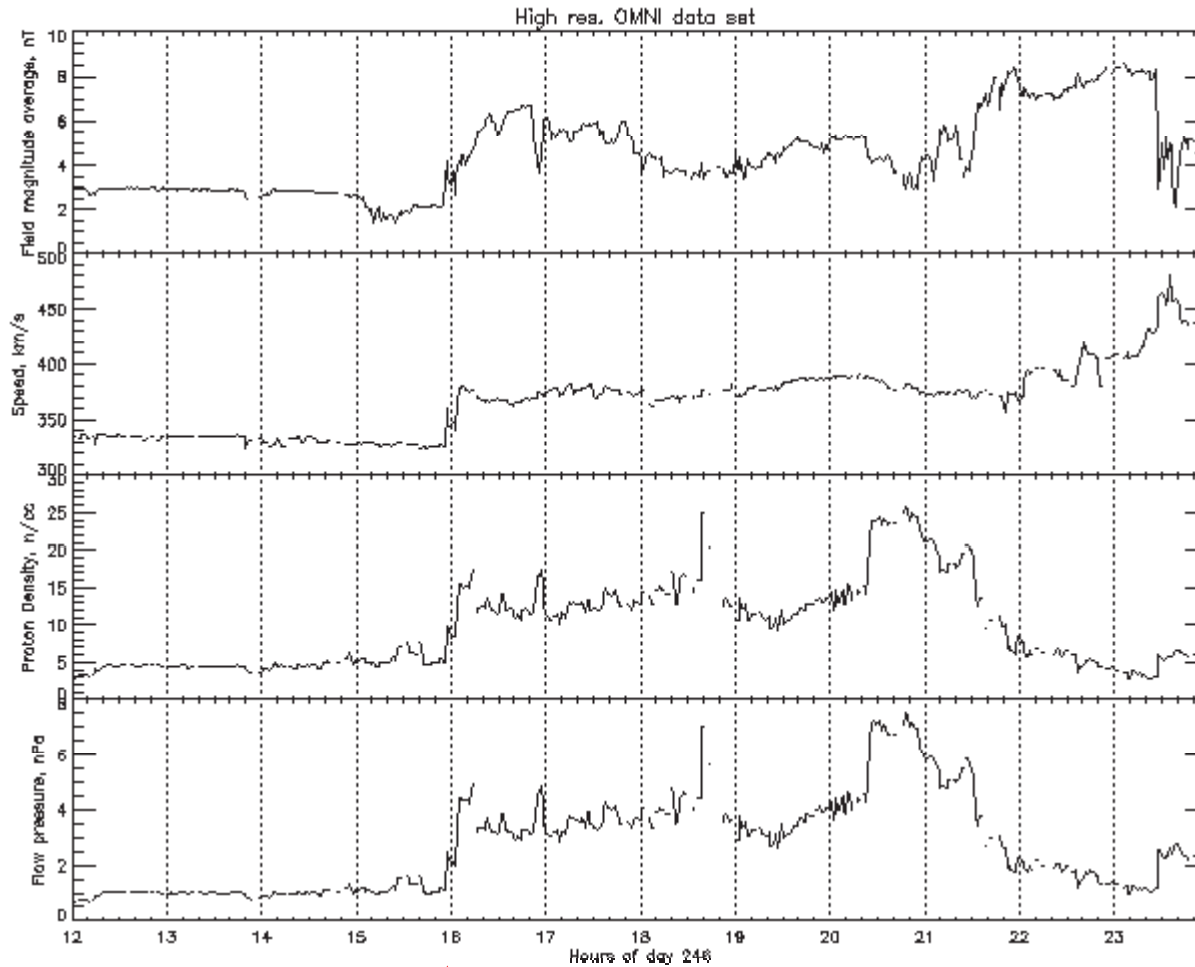
Case-5 Sep 3, 2009

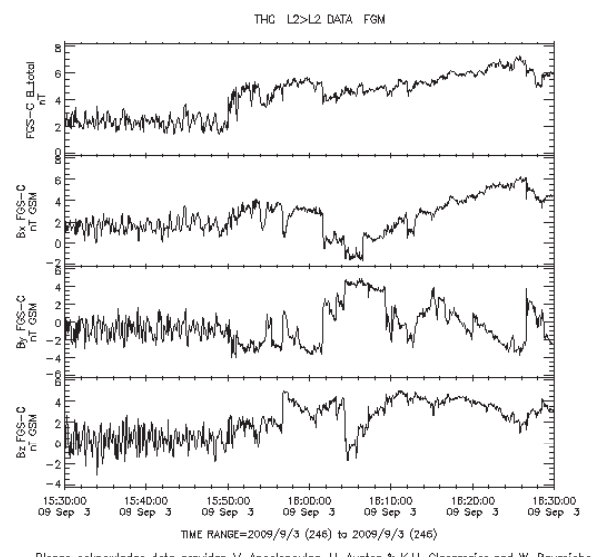
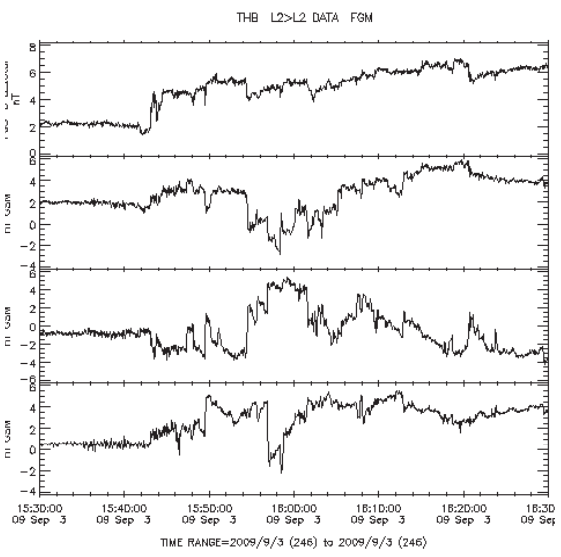
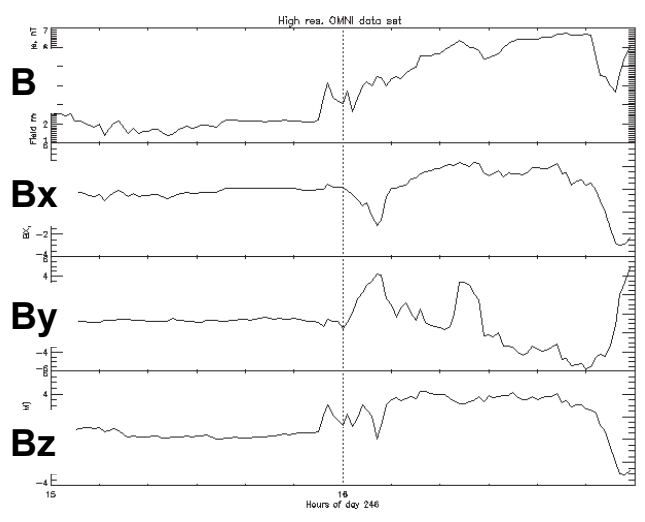
Btot

Vsw

Nsw

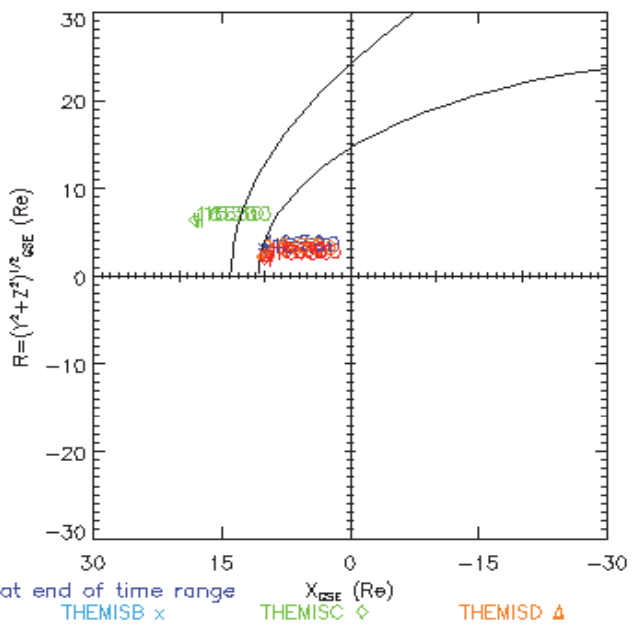
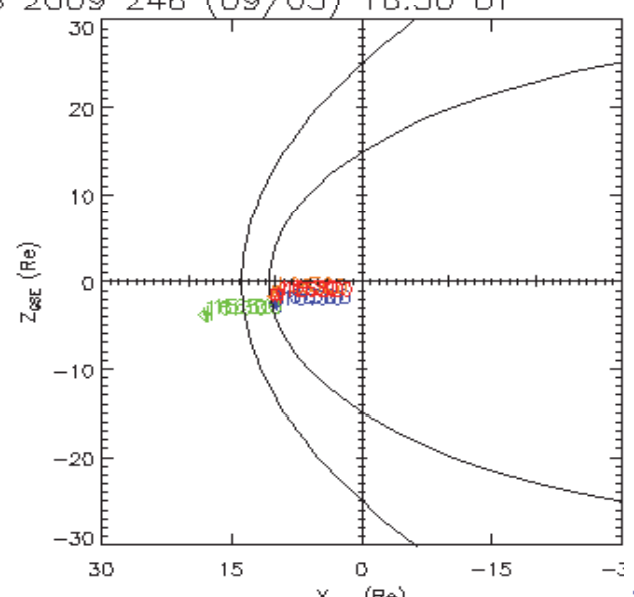
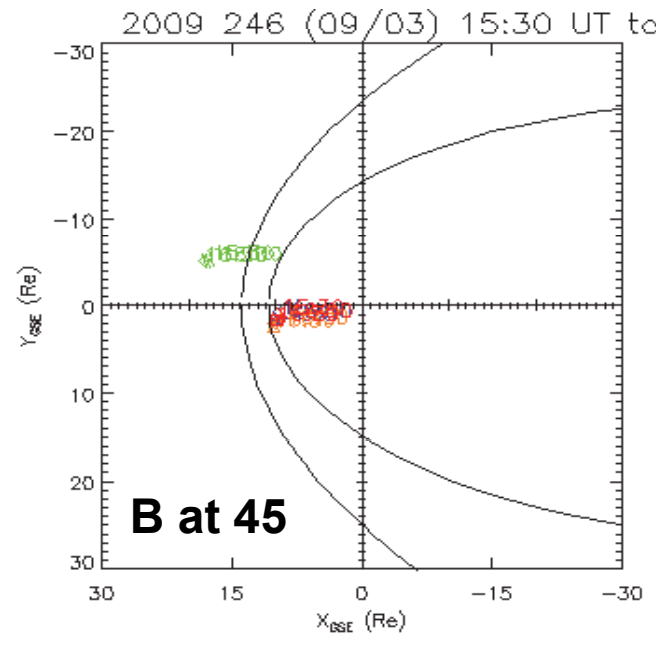
Pdyn



B**C**

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TA

Bx

By

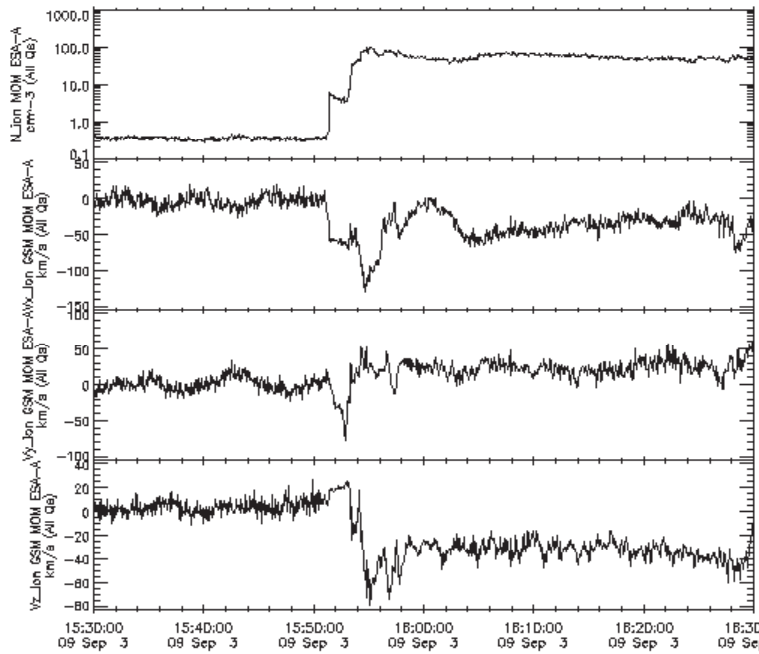
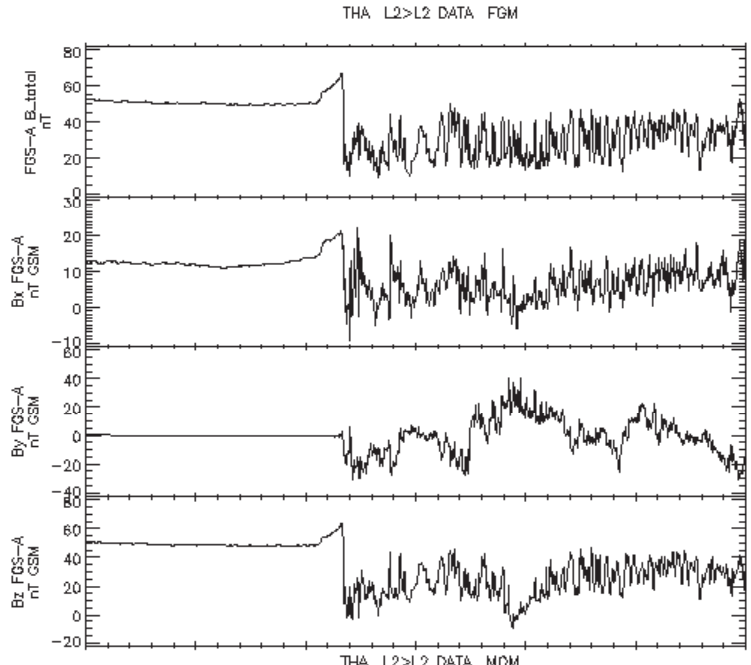
Bz

n

Vx

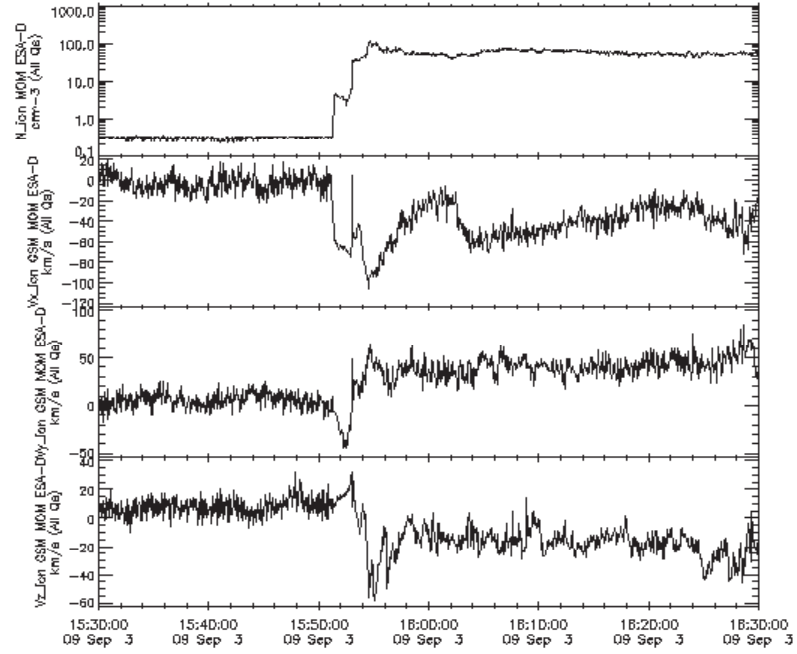
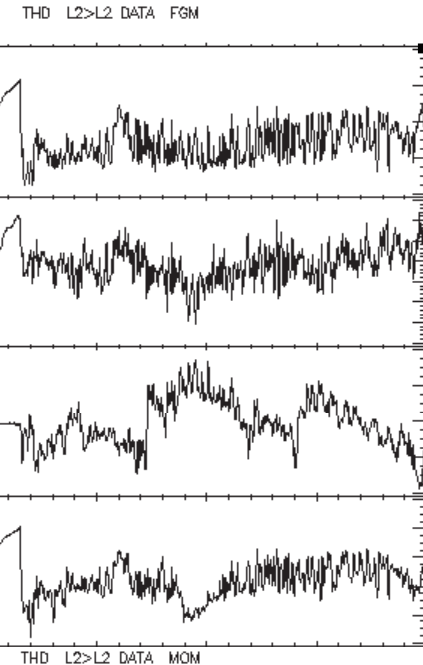
Vy

Vz



TIME RANGE=2009/9/3 (246) to 2009/9/3 (246)

Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDAWeb when using these data.
Generated by CDAWeb on Sun Sep 11 22:39:26 2011



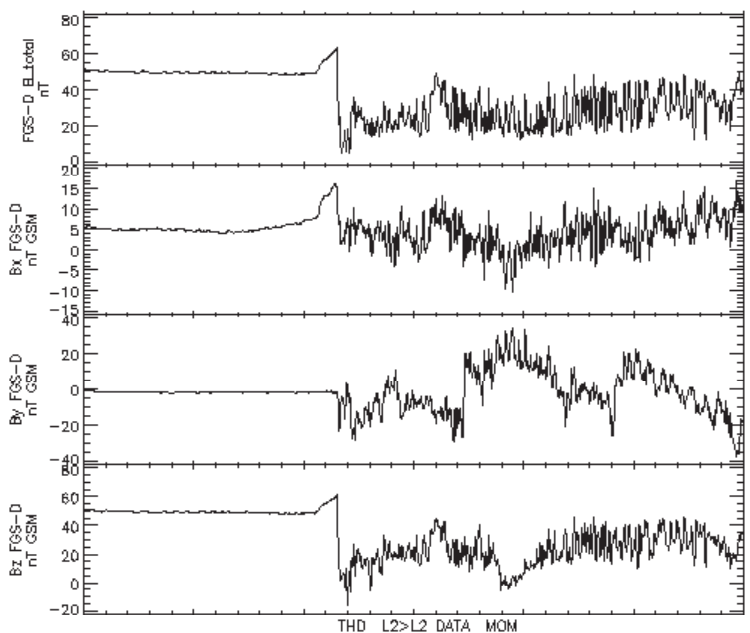
TIME RANGE=2009/9/3 (246) to 2009/9/3 (246)

Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDAWeb when using these data.
Generated by CDAWeb on Sun Sep 11 22:39:26 2011

TD

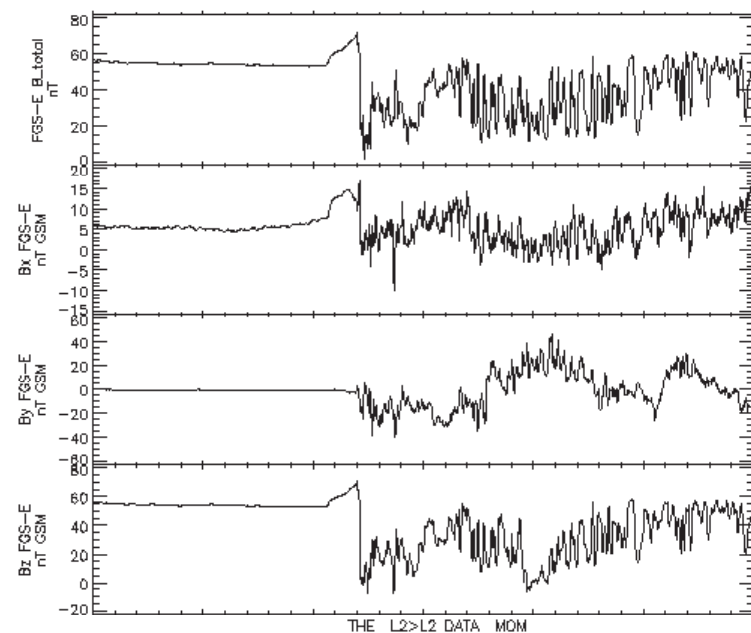
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THD L2>L2 DATA FGM

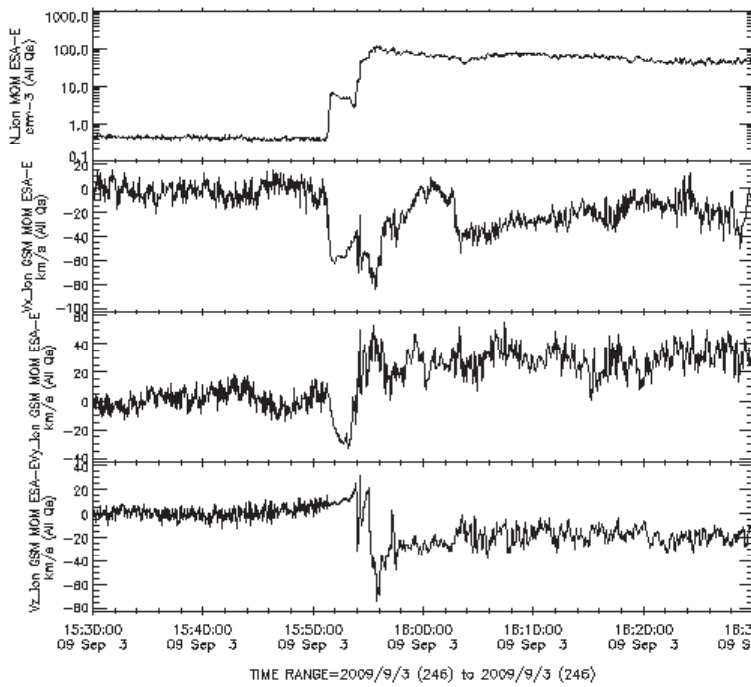
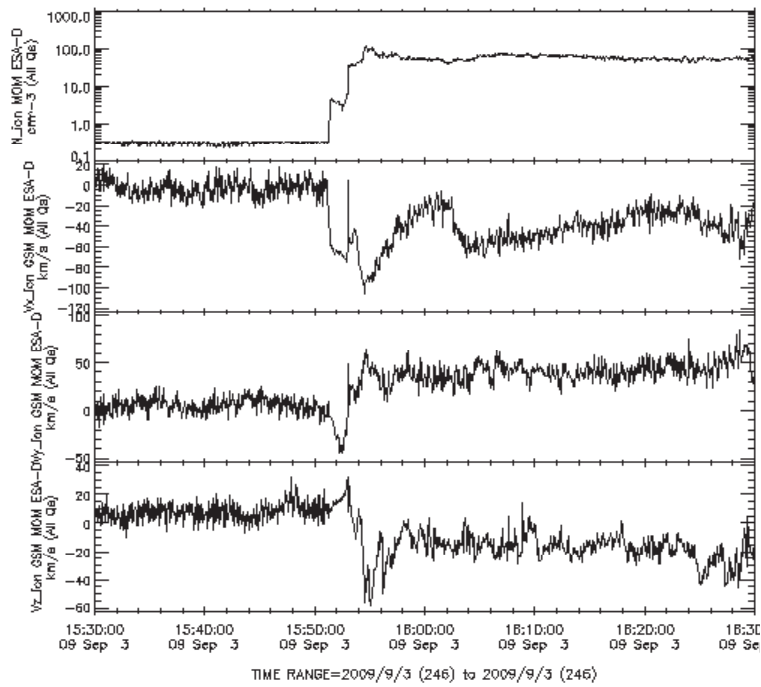


TE

THE L2>L2 DATA FGM



n



Vx

Vy

Vz

Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden
at UCB, NASA NASS-02099 and CDAWeb when using these data.
Generated by CDAWeb on Sun Sep 11 22:39:26 2011

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at UCB, NASA NASS-02099 and CDAWeb when using these data.
Generated by CDAWeb on Sun Sep 11 22:39:26 2011

TD

Bx

By

Bz

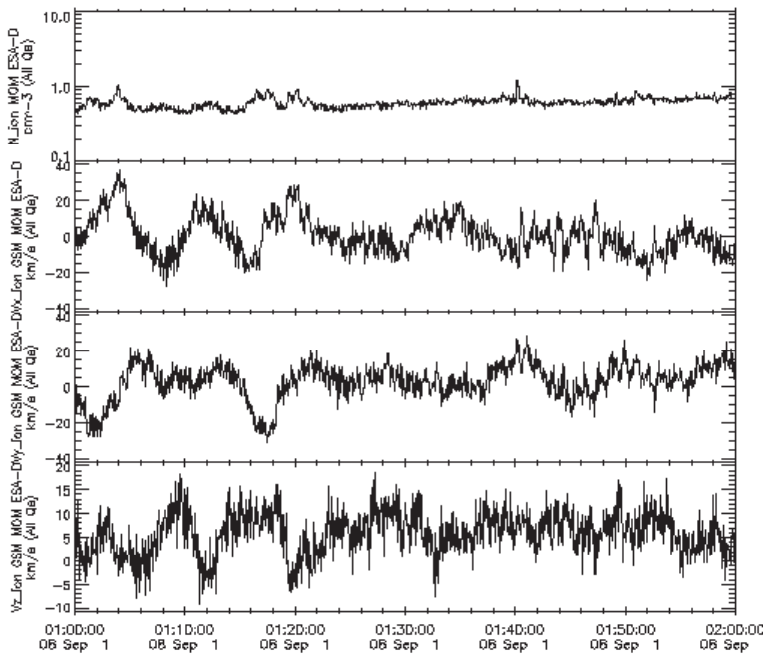
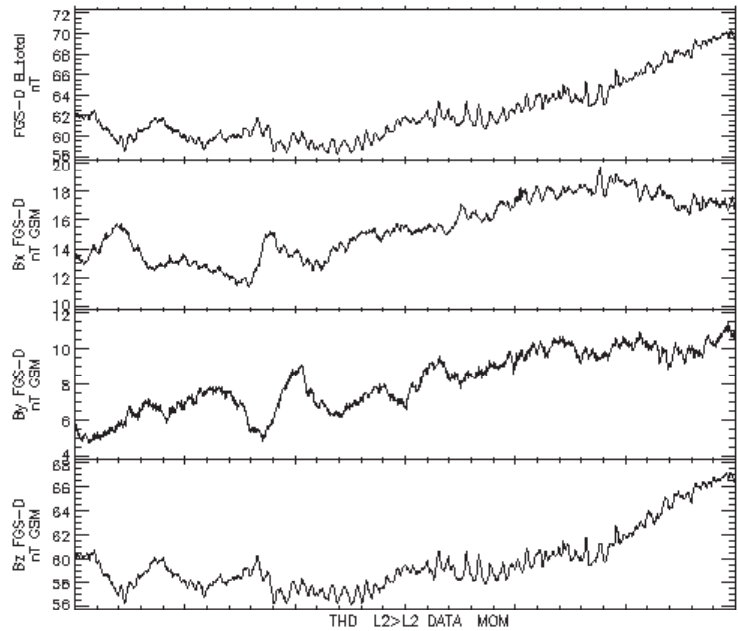
n

Vx

Vy

Vz

THD L2>L2 DATA FGM

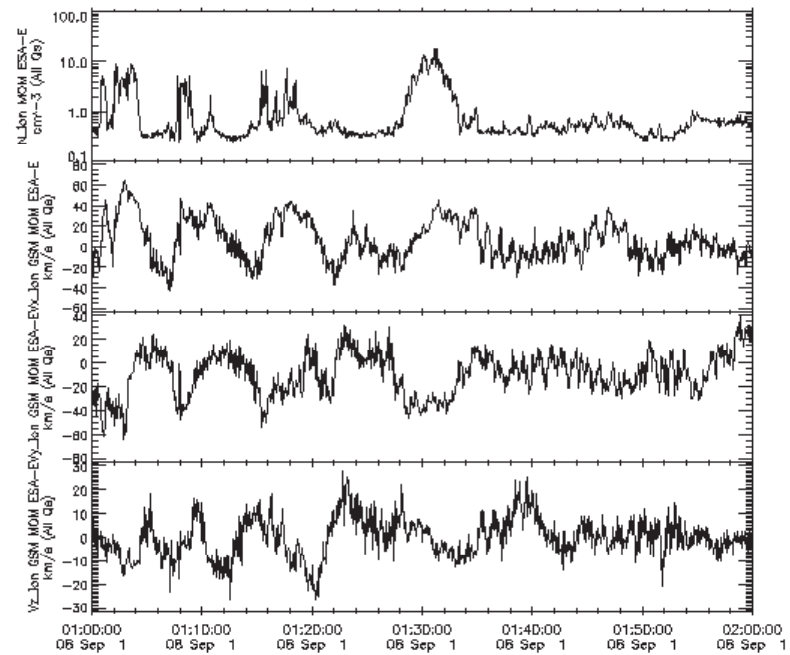
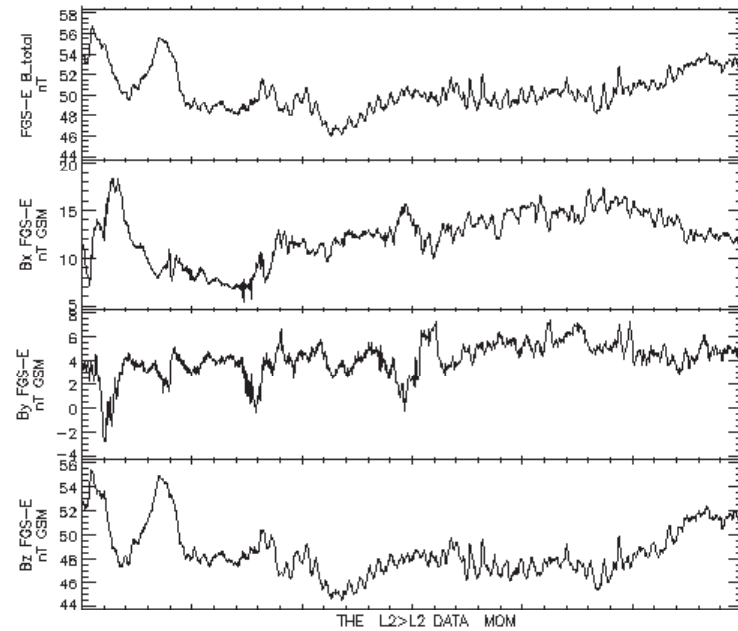


TIME RANGE=2008/9/1 (245) to 2008/9/1 (245)

Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDWeb when using these data.
Generated by CDWeb on Tue Sep 13 10:19:39 2011

THE L2>L2 DATA FGM

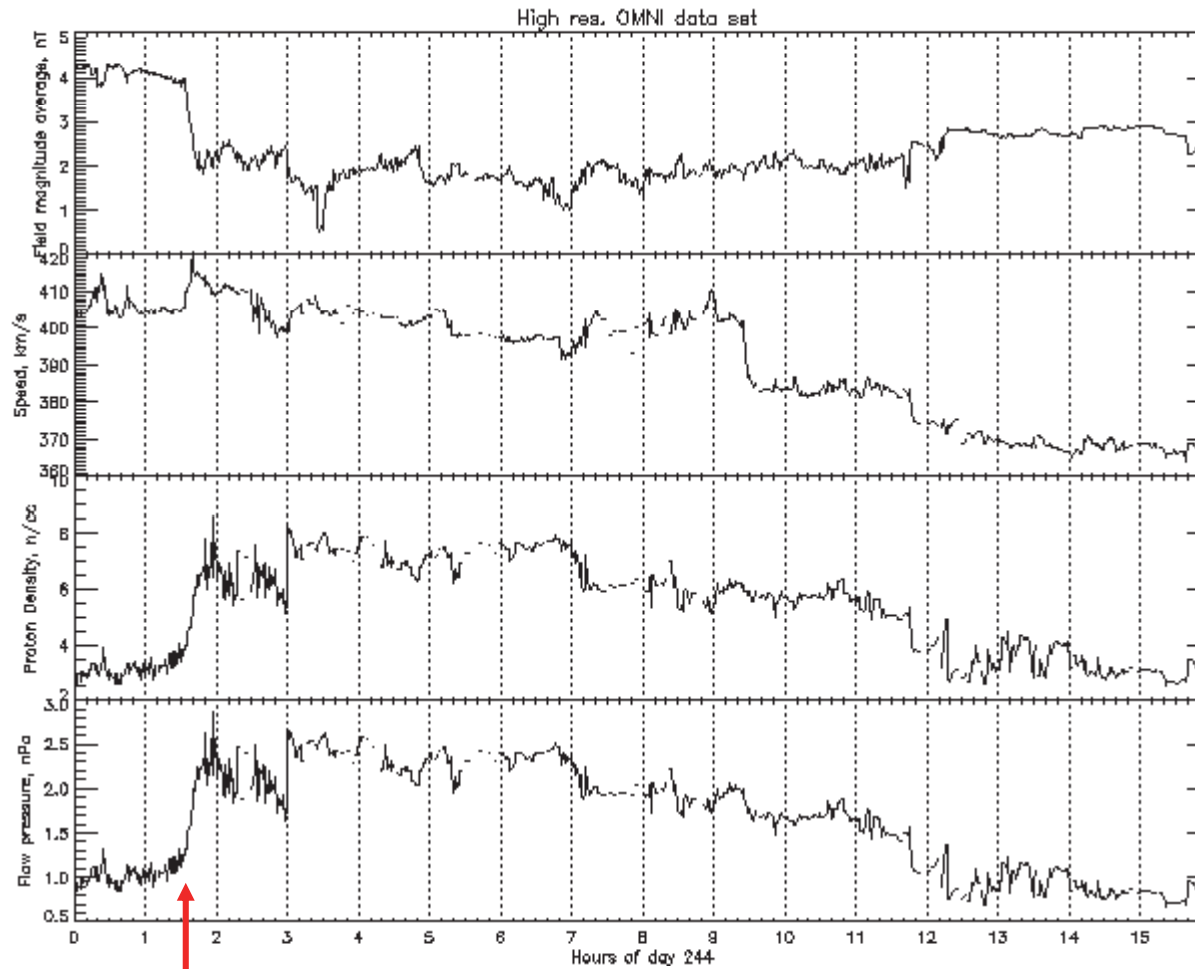
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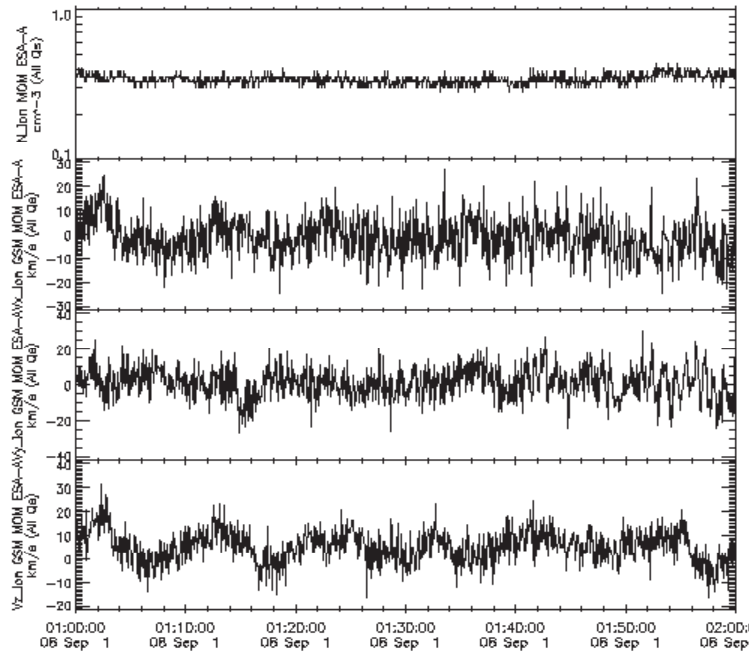
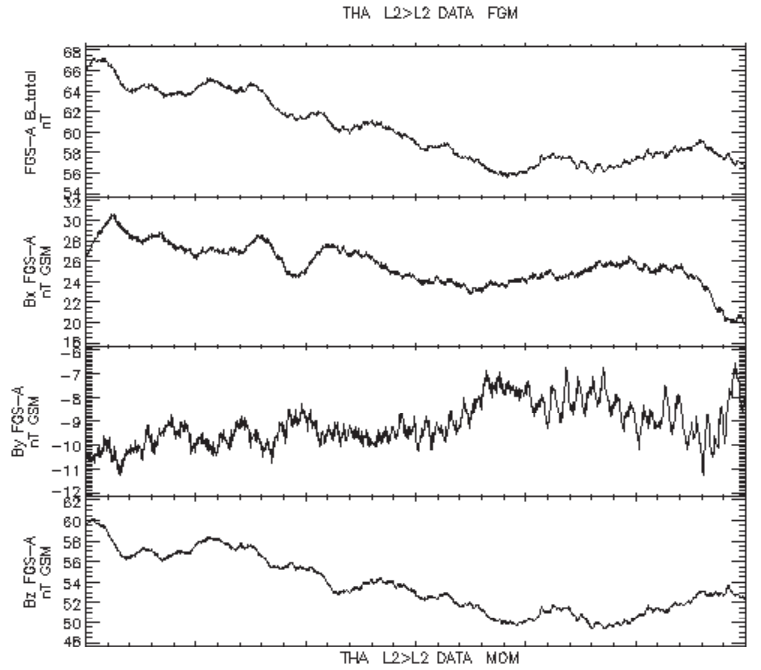
TIME RANGE=2008/9/1 (245) to 2008/9/1 (245)

Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDWeb when using these data.
Generated by CDWeb on Tue Sep 13 10:19:39 2011

Case-6: Sep 1, 2009

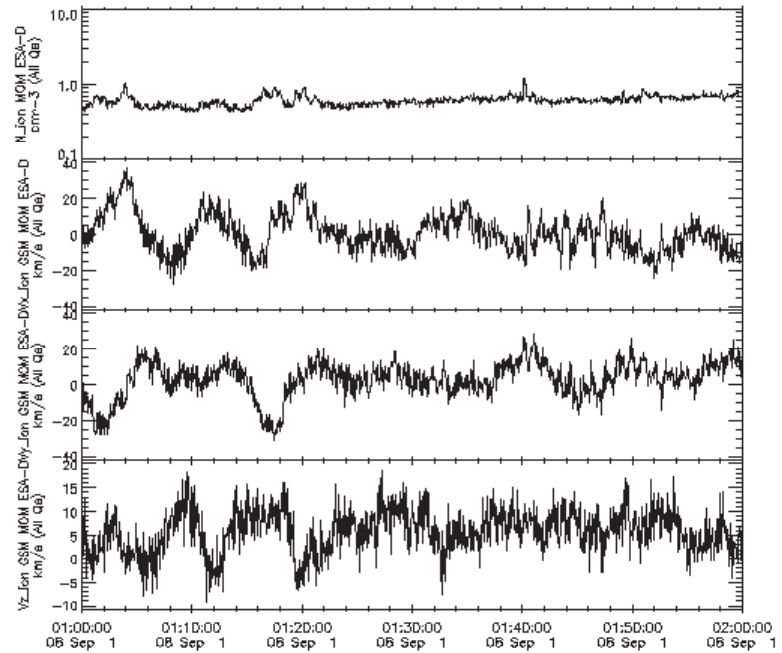
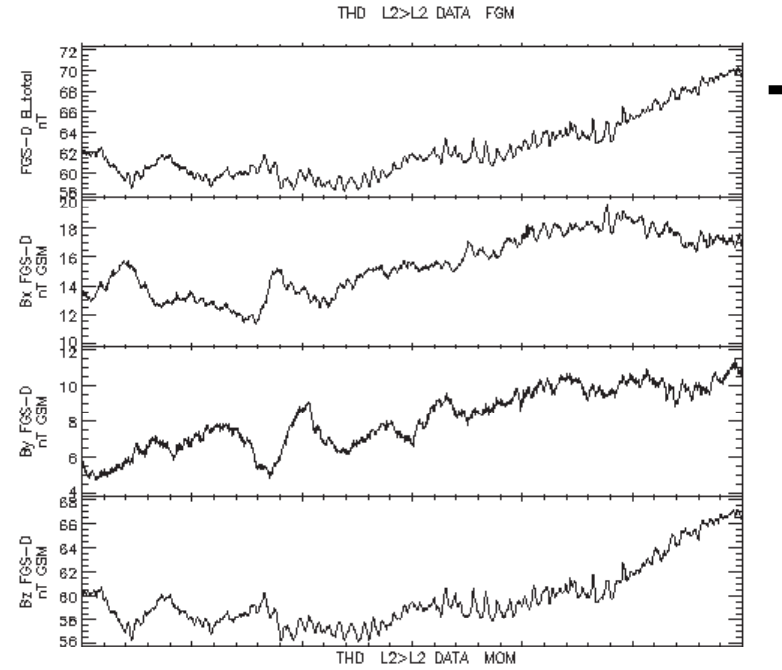


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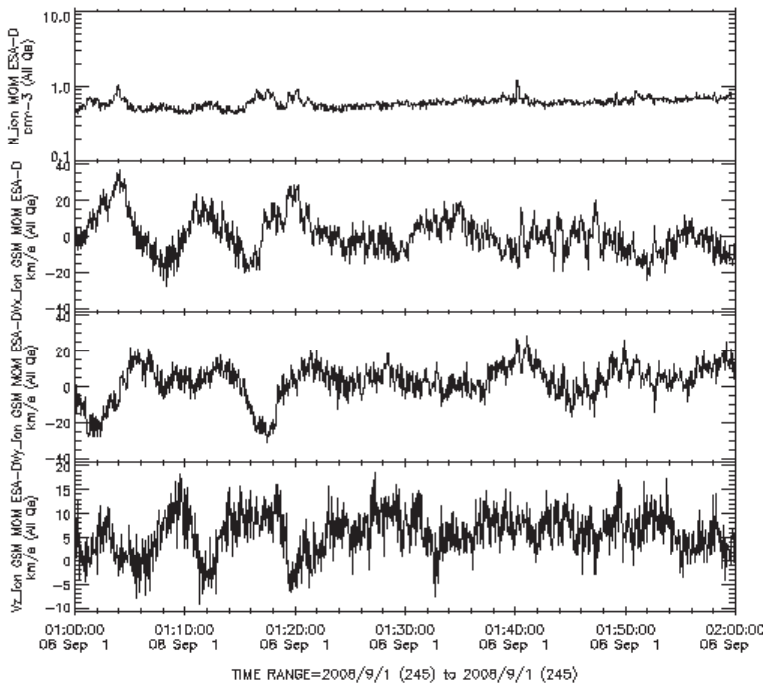
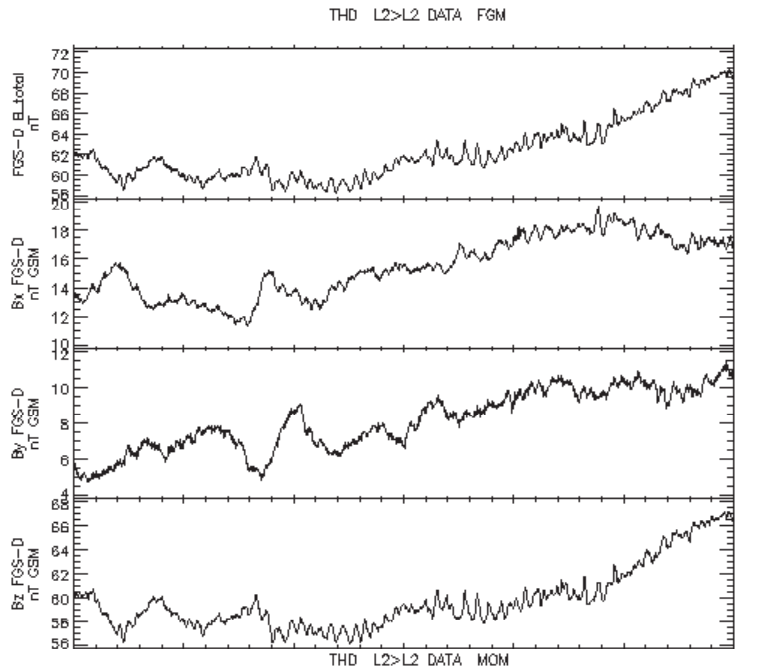
Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NAS5-02099 and CDAWeb when using these data.
Generated by CDAWeb on Tue Sep 13 10:19:38 2011

TD



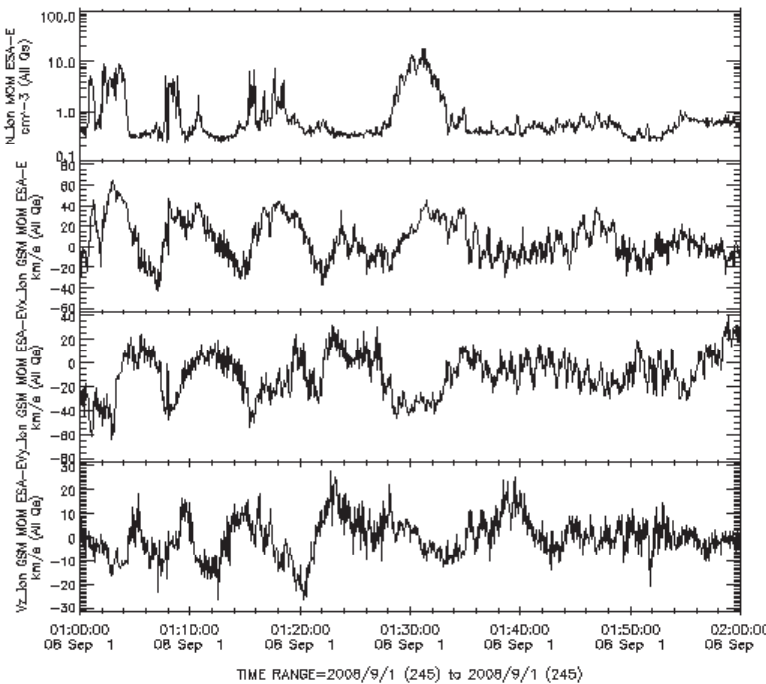
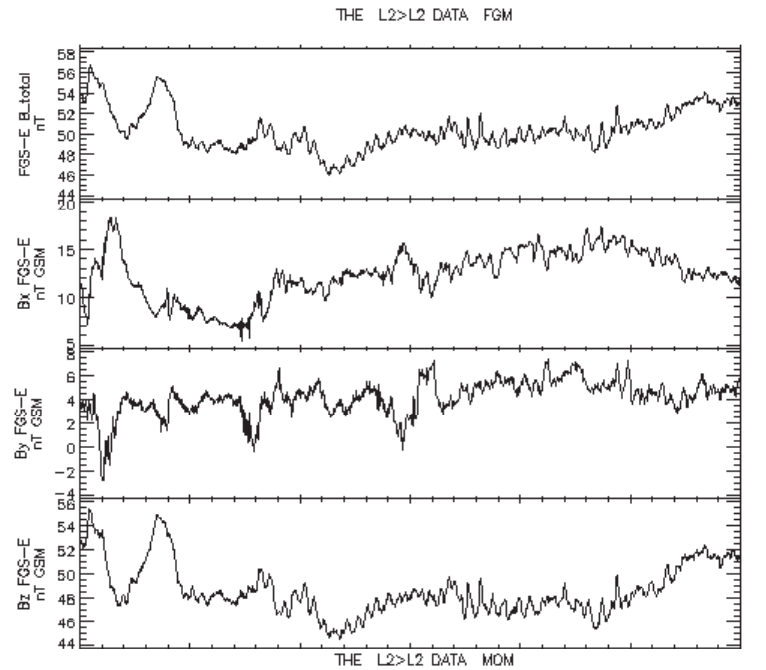
Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NAS5-02099 and CDAWeb when using these data.
Generated by CDAWeb on Tue Sep 13 10:19:39 2011

TD



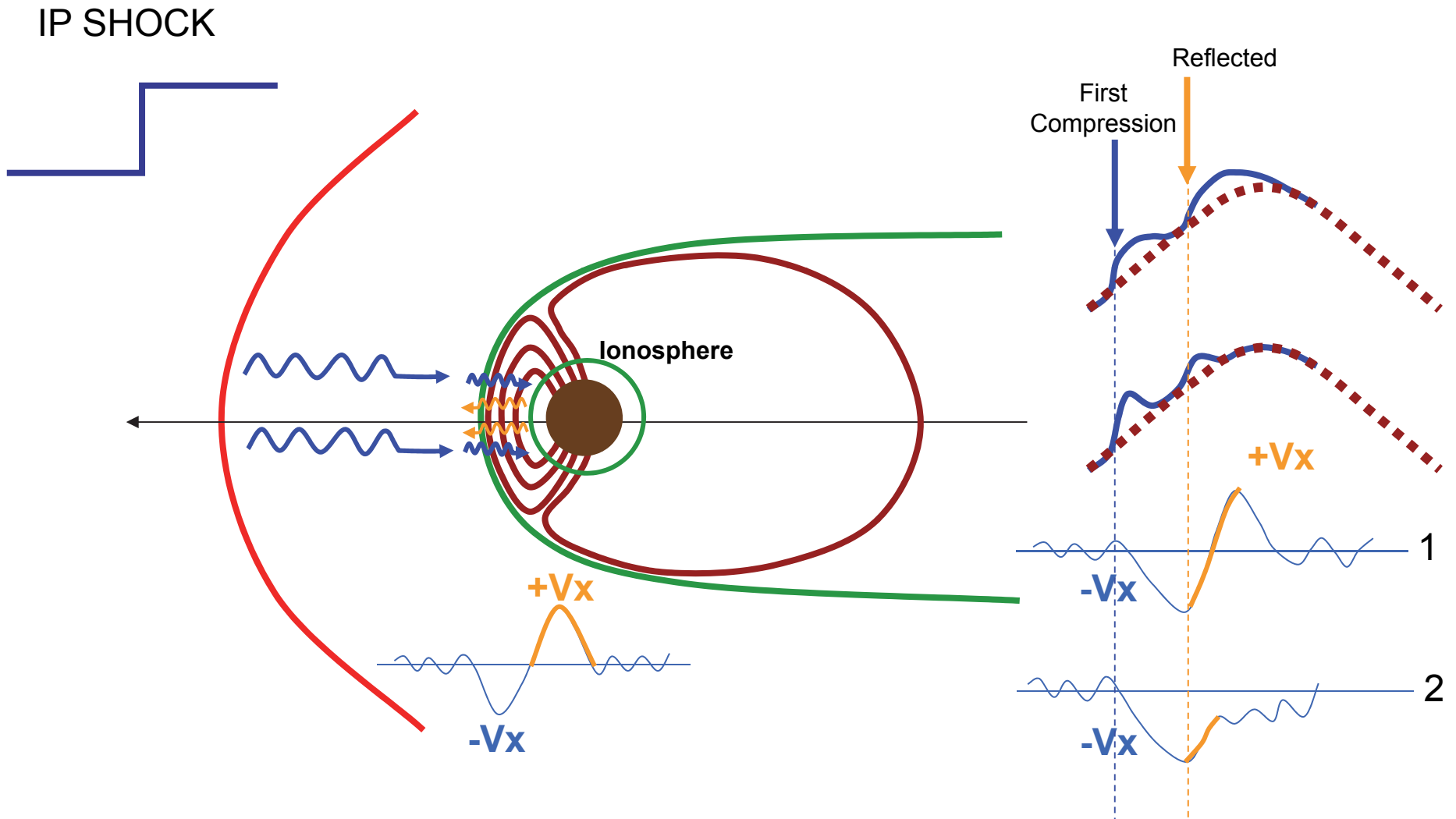
Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDAWeb when using these data.
Generated by CDAWeb on Tue Sep 13 10:19:39 2011

TE



Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDAWeb when using these data.
Generated by CDAWeb on Tue Sep 13 10:19:39 2011

Concept of Reflected Wave



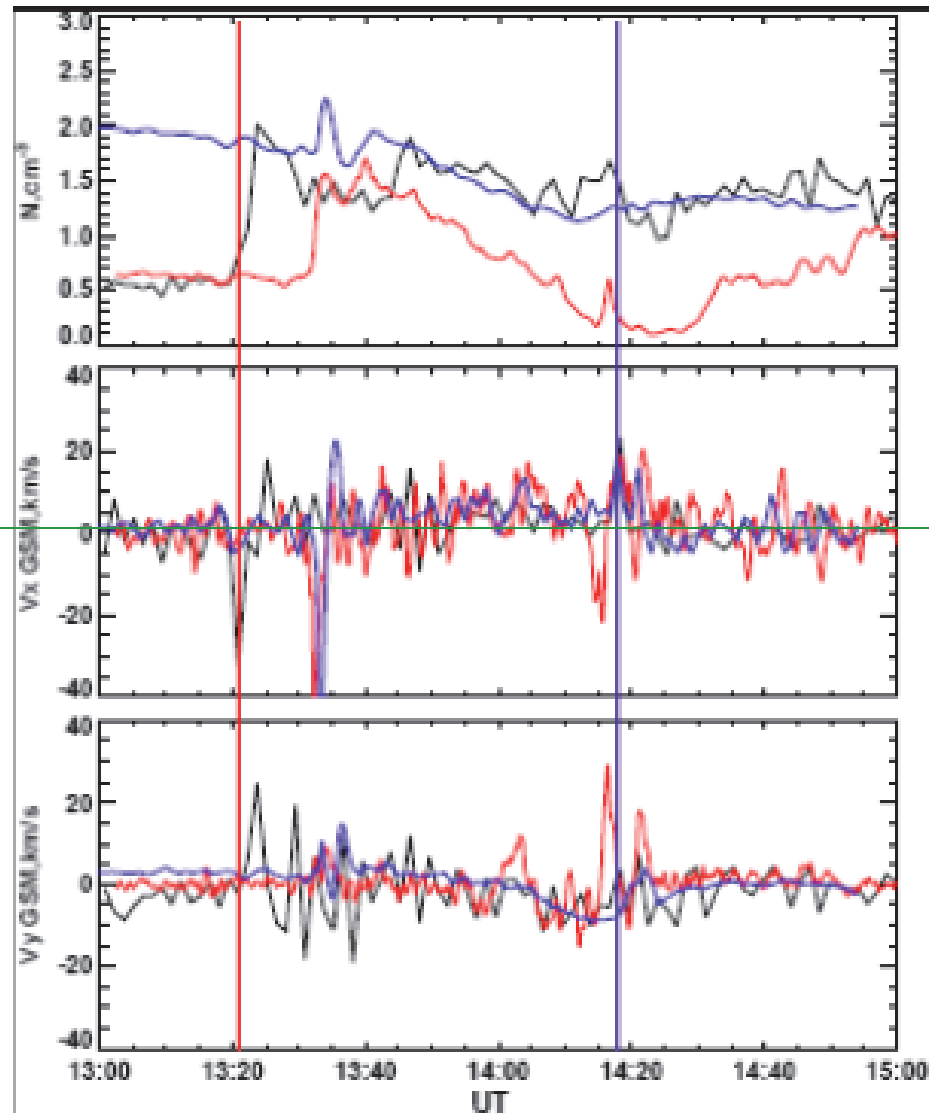
Concept of reflected wave

- Ridley 2006
 - Results from global MHD model, BATSRUS
 - Simulating interaction between strong IP shock and the magnetosphere
 - Transmitted through msheath and msphere
 - Reflected from inner numerical boundary
 - Continue moving outward towards mpause
 - Assumed reflection boundary is plasmopause
 - Transmitted IP shock propagated around the flanks of the mpshere toward mtail
 - Dayside bow shock moved inward until it interacted with the reflected shock

Concept of reflected wave

- Lee and hudson (2001)
- 3-D dipole model for magnetosphere
- To study Similar problem
- Propagation of a sudden impulse associated with IP shock inside the msphere
- To simulate th eresponse they invoked an abrupt electric field variation at the outer boundary of their model
 - Their results indicate that:
 - Most of impulse energy penetrates the plasmopause to excite strong low frequency pulsations in th eplasmaosphere
 - Small portion of th einitial impulse however is reflected from the plasmopause and returns to the outer boundary

Concept of reflected wave



LANL ion moment Measurements (black lines)

Red: batsrus

Blue: LFM

Samsonov et al, 2007

Summary

- We found 40 clear box type IP shock events during 2008-2010.
 - In response to the IP shock magnetosphere oscillates
 - The motion is easier in the y-direction, least in the z-direction
 - Variations last 10-40 min (recovery)
 - Speeds both in x and y vary between 20-100 km/sec
 - Clear Reflected wave among the events we found is not seen → needs more work
 - Boundary Normal coordinate analysis is under way to understand the motion of the magnetosphere better

Welcome

- Suggestions and comments

Thank you and

- THEMIS mission
- Nasa's guest program

Summary

- Dayside response of the magnetosphere
- We found 40 clear box type IP shock events during 2008-2010.
- We excluded
 - multiple shocks and
 - weak shocks where jump in pressure less than 2 factor of background
- Depends on the location of s/c
 - 45° of Sun-Earth line

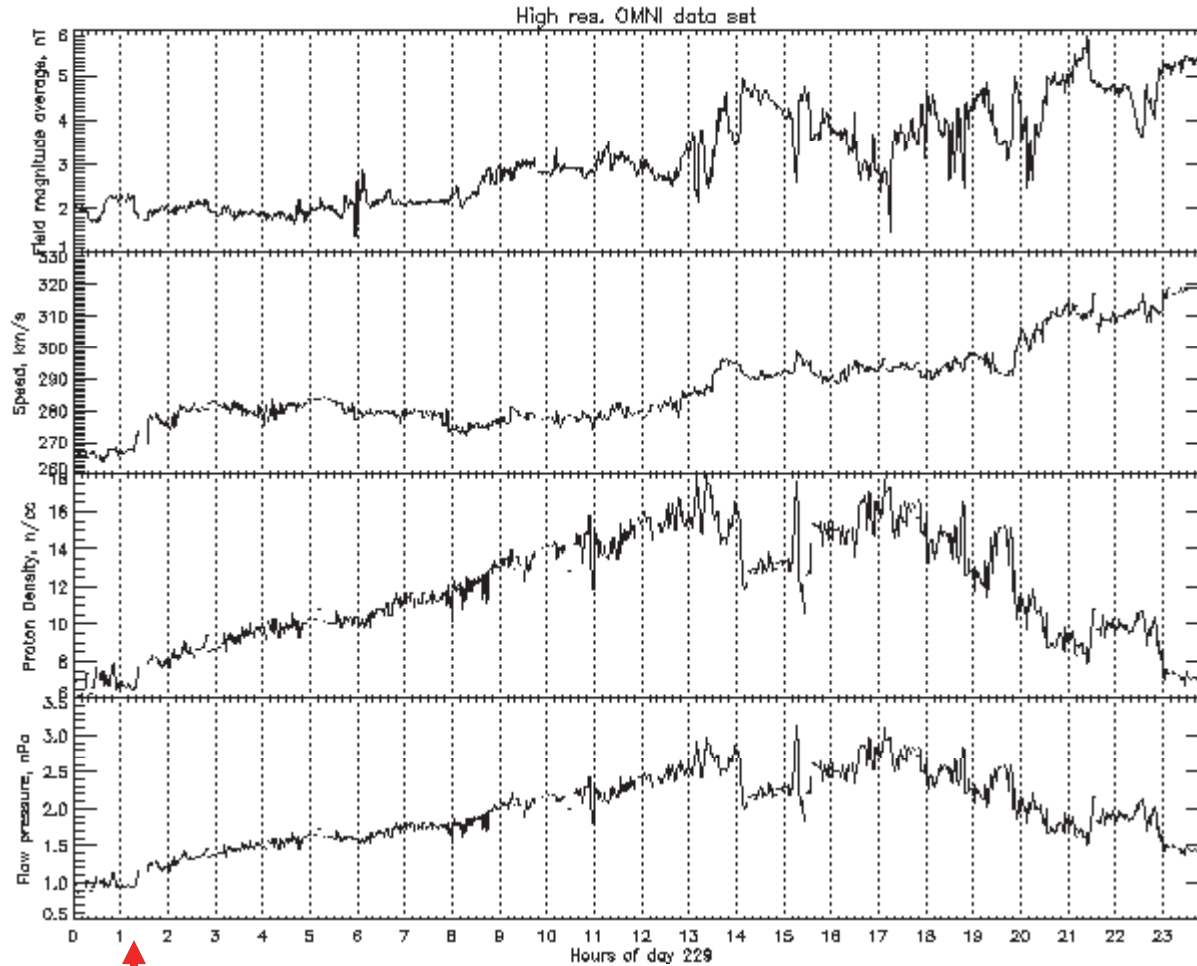
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- The motion is easier in the y-direction, least in the z- direction
- Variations last 10-40 min (recovery)
- Speeds both in x and y vary between 20-100 km/sec
- Clear Reflected wave among the events we found is not seen → needs more work
- Boundary Normal coordinate analysis is under way to understand the motion of the magnetosphere better
- -----
- In some cases, the IP signal is weak but still drives magnetosphere at the same rates
- In some cases, IP is strong but the response is weak.
 - Response in term of magnitude of the periodic oscillations

- Study is underway with
 - Electric field measurements
 - GOES measurements
 - Ground magnetometer
- Any of my events (50), no Cluster and Geotail measurements on the dayside are available

Thank you and

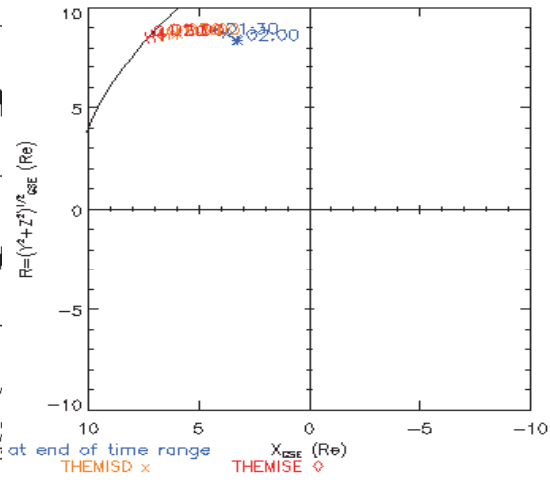
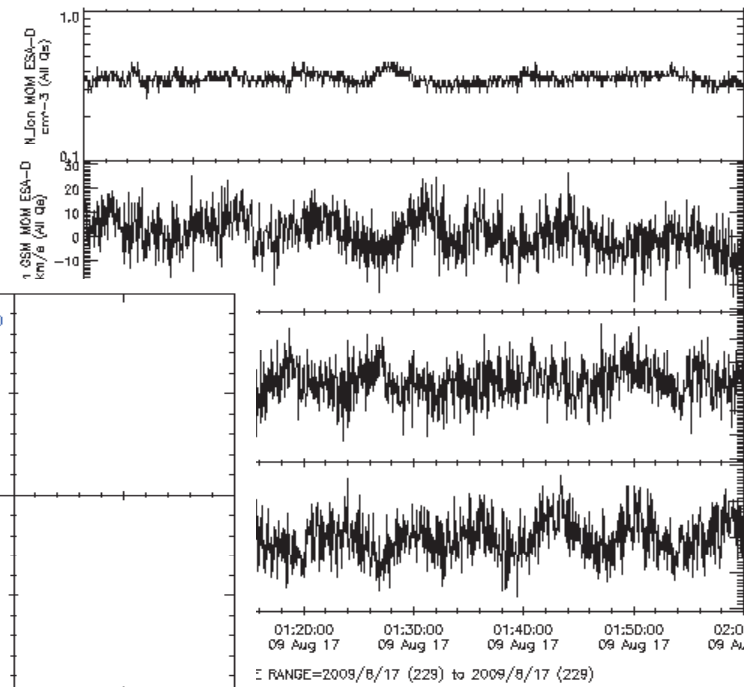
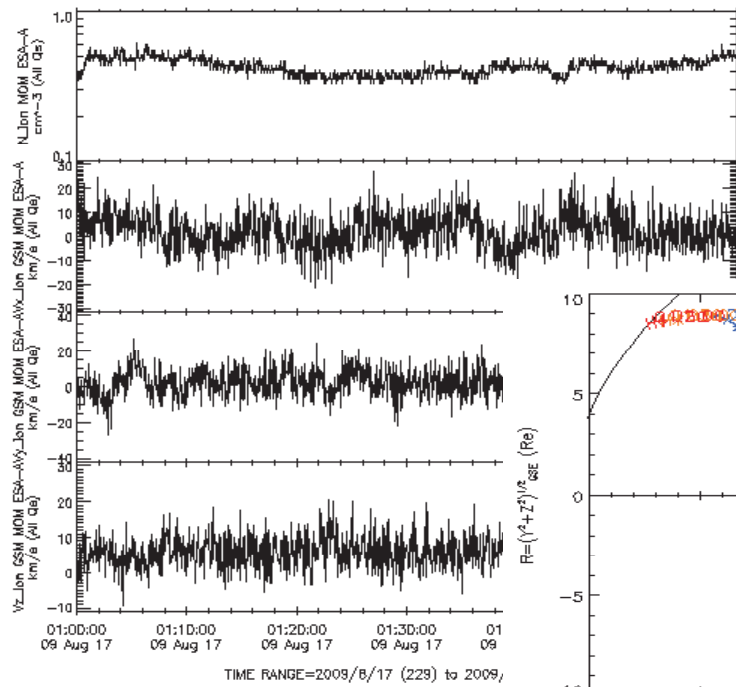
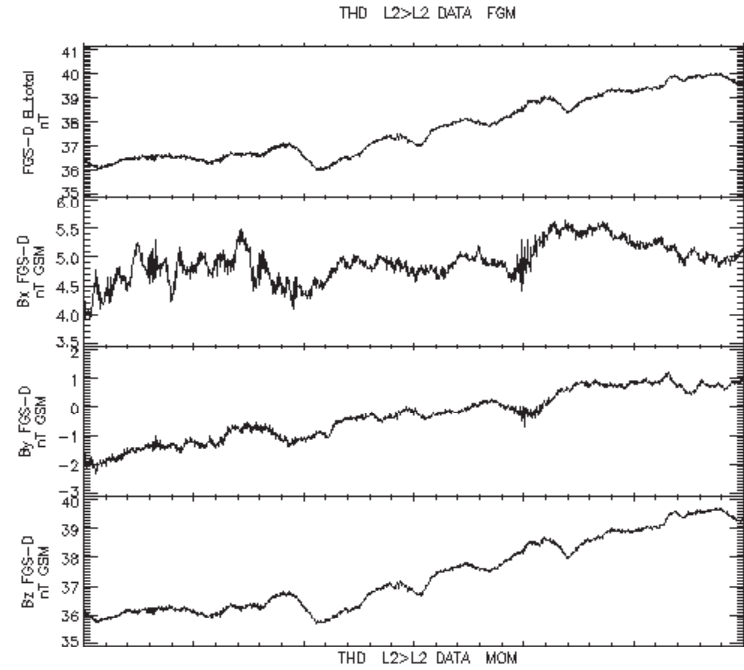
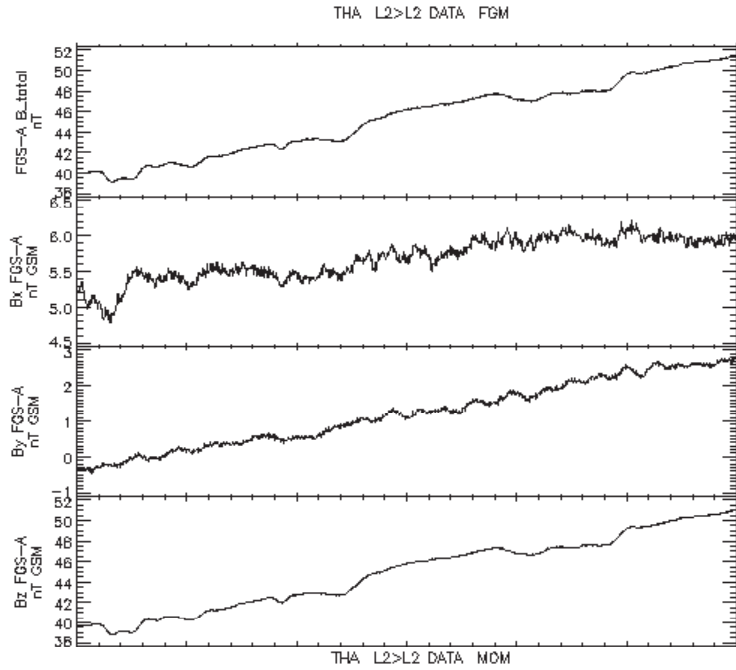
- THEMIS mission
- Nasa's guest program

Case 7: Aug 17, 2009



Weak Shock

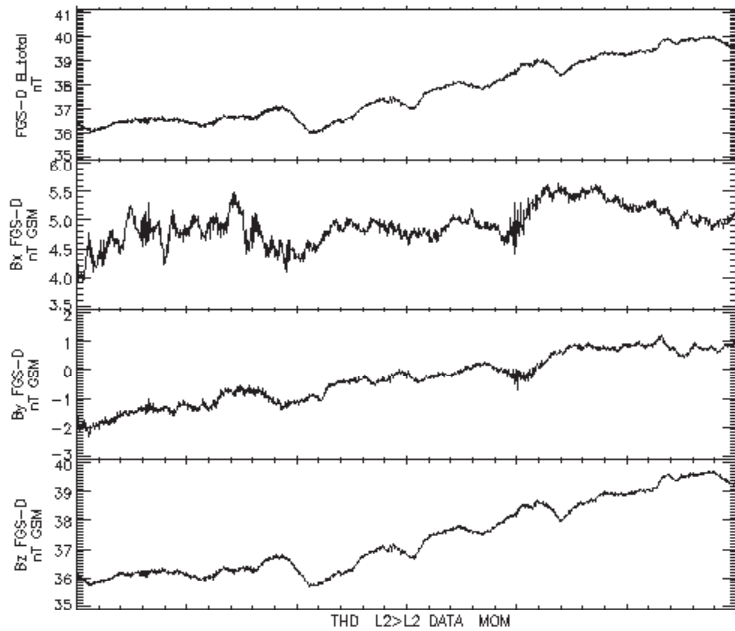
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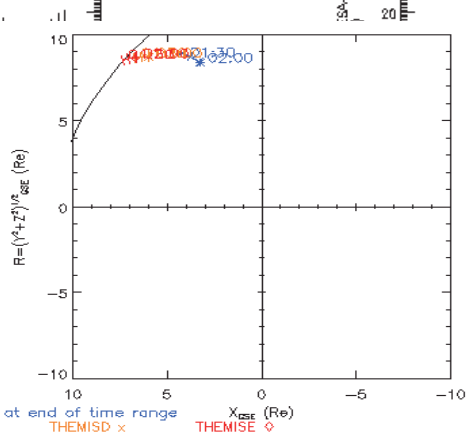
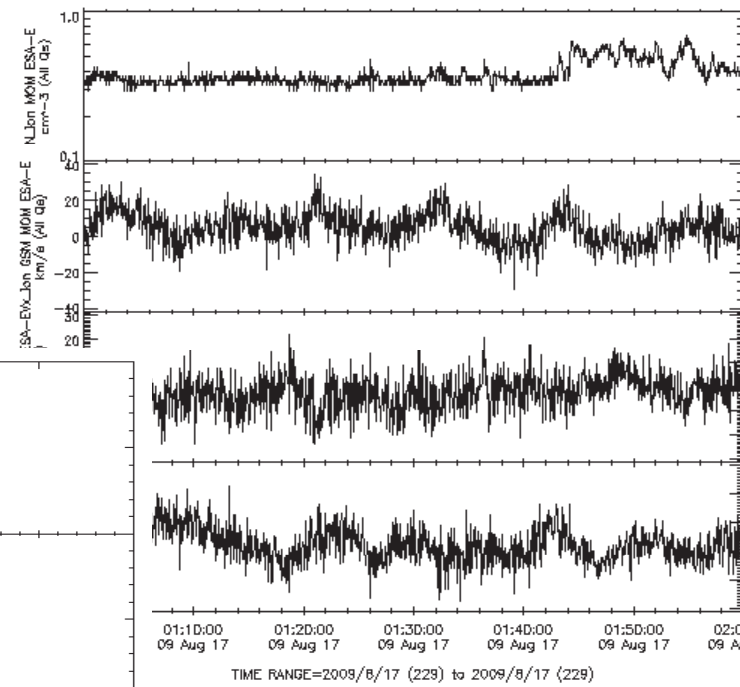
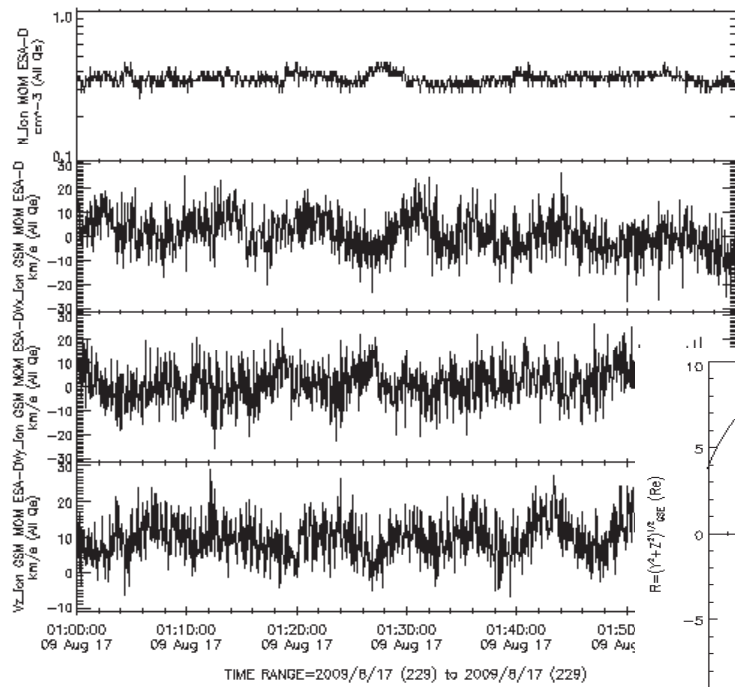
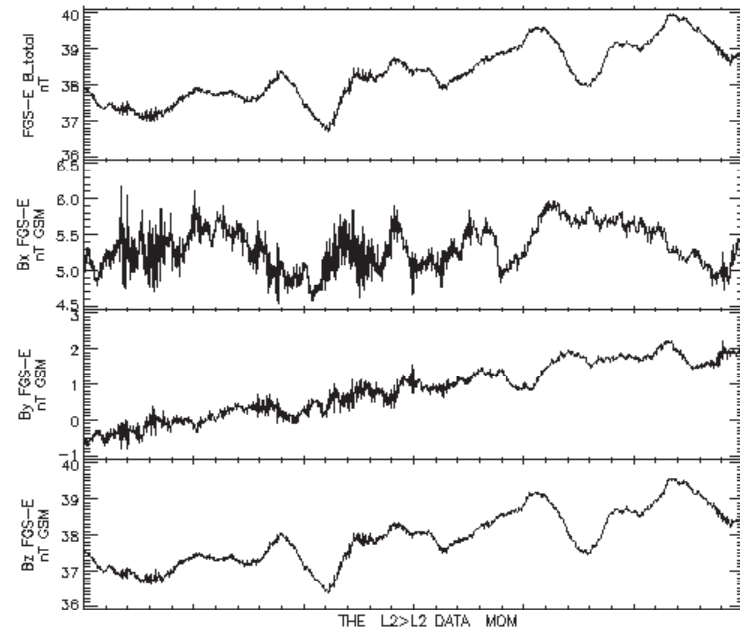
Please acknowledge data provider, V. Angelopoulos, at UCB, NASA NASS-02099 and CDAWeb when using these data. Generated by CDAWeb on Tue Sep 13 11:47:46 2011

Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDAWeb when using these data. Generated by CDAWeb on Tue Sep 13 11:47:46 2011

THD L2>L2 DATA FGM



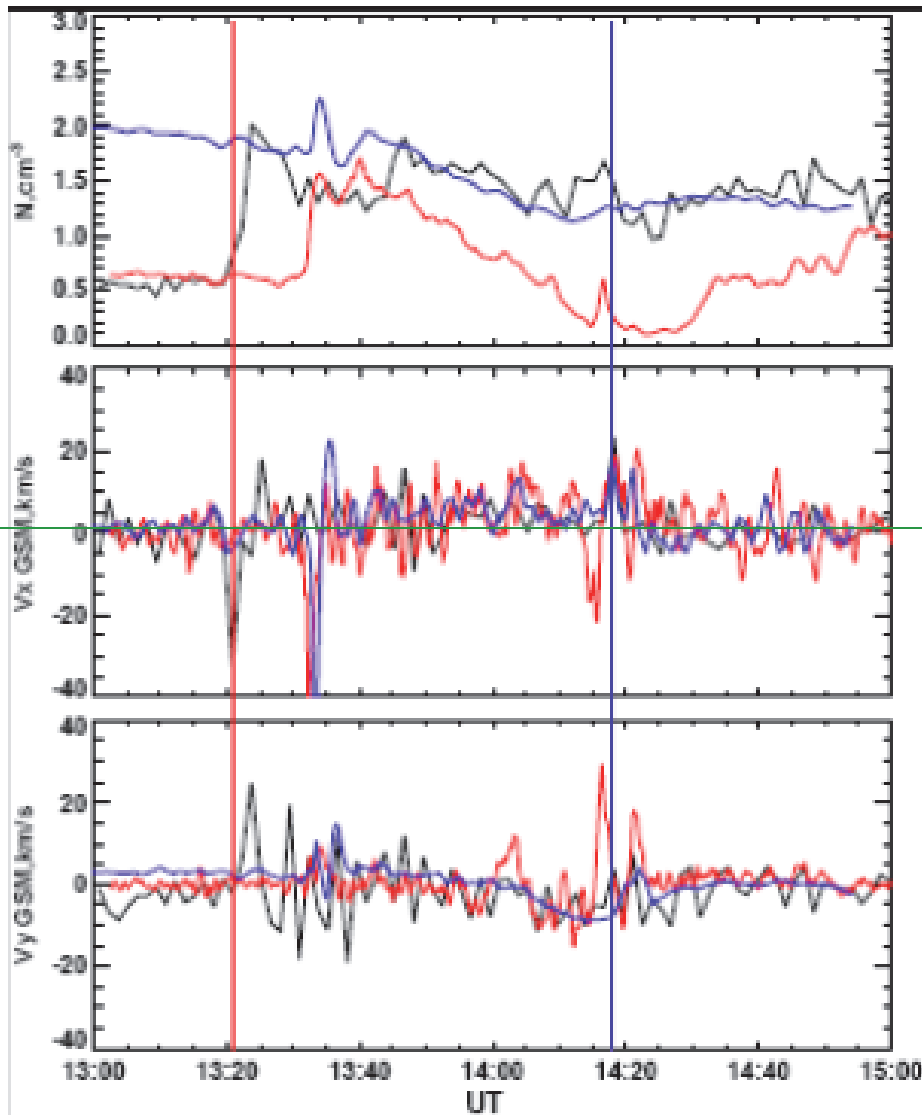
THE L2>L2 DATA FGM



Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDWeb when using these data.
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Please acknowledge data provider, V. Angelopoulos, C.W. Carlson & J. McFadden at UCB, NASA NASS-02099 and CDWeb when using these data.
Generated by CDWeb on Tue Sep 13 11:47:46 2011

Concept of reflected wave

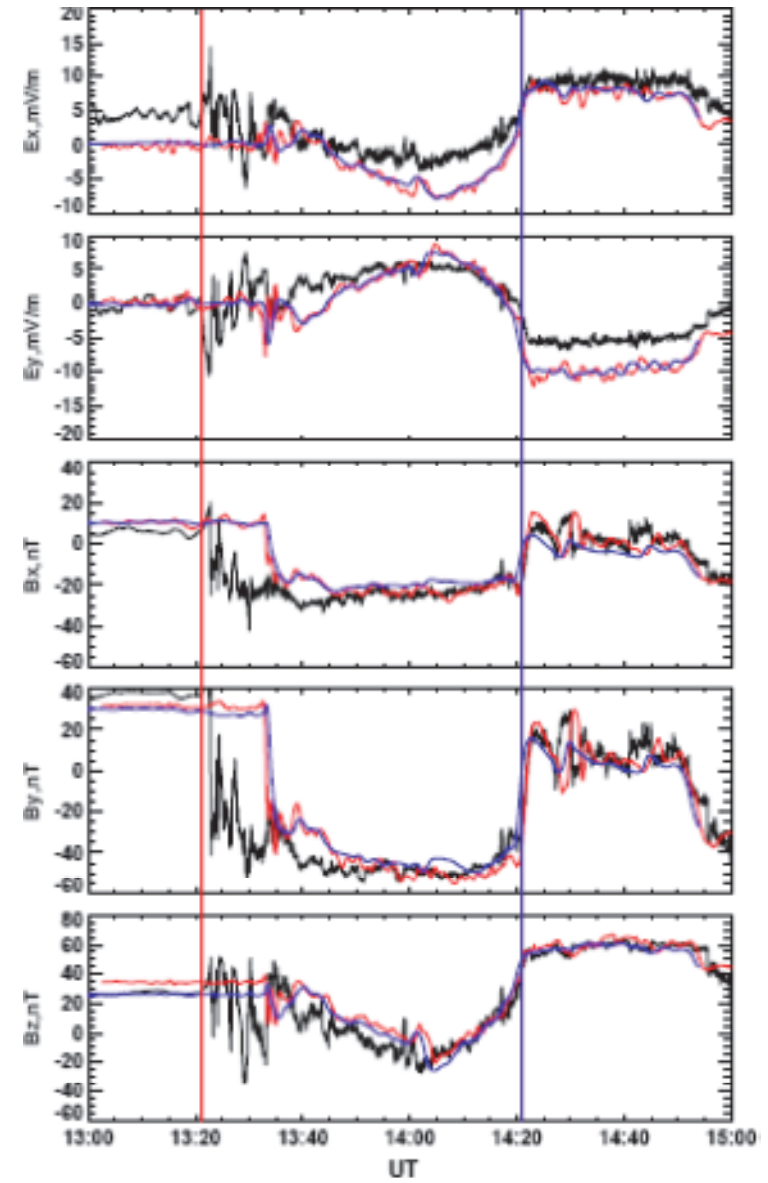


LANL ion moment Measurements (black lines)

Red: batsrus

Blue: LFM

Samsonov et al, 2007



Geotail measurements of E field and B field